

Decoding the Standard Model with Flavour Physics

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Outline:

1. The Standard Model and the Flavour Problem
2. The V_{cb} puzzle
3. Outlook and prospects on BSM

Motivation

Despite the SM successes,
there are open problems:

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Hierarchy problem

dark matter/dark energy

flavour hierarchies

neutrino masses

gravity

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SM(EFT)

Λ_{EW}

Energy

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UV theory

SM(EFT)

Λ_{UV}

Λ_{EW}

Energy

Motivation

Despite the SM successes,
there are open problems:

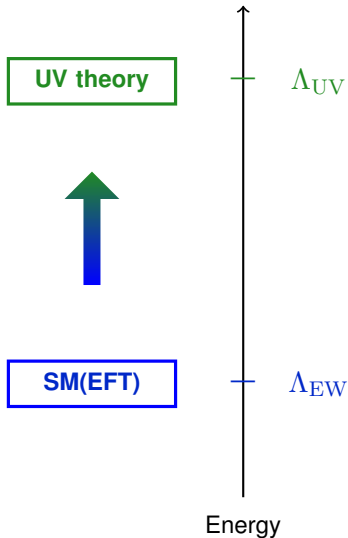
Hierarchy problem

dark matter/dark energy

flavour hierarchies

neutrino masses

gravity



The (two) flavour problems

1. **The SM flavour problem:** The measured Yukawa pattern doesn't seem accidental

⇒ Is there any deeper reason for that?

2. **The NP flavour problem:** If we regard the SM as an EFT valid below a certain energy cutoff Λ , why don't we see any deviations in flavour changing processes?

⇒ Which is the flavour structure of BSM physics?

The SM flavour problem

$$\mathcal{L}_{\text{Yukawa}} \supset Y_u^{ij} \bar{Q}_L^i H u_R^j$$

$$Y_u \sim y_t \begin{pmatrix} \text{light green circle} & \text{light green circle} & \text{dark green circle with } 0.003 \\ & \text{dark green circle} & \text{dark green circle with } 0.04 \\ & & 1 \end{pmatrix}$$

The SM flavour problem

$$\mathcal{L}_{\text{Yukawa}} \supset Y_u^{ij} \bar{Q}_L^i H u_R^j$$

$$Y_u \sim y_t \begin{pmatrix} & & \\ & & \\ & & 1 \end{pmatrix}$$

Exact $U(2)^n$ limit

The SM flavour problem

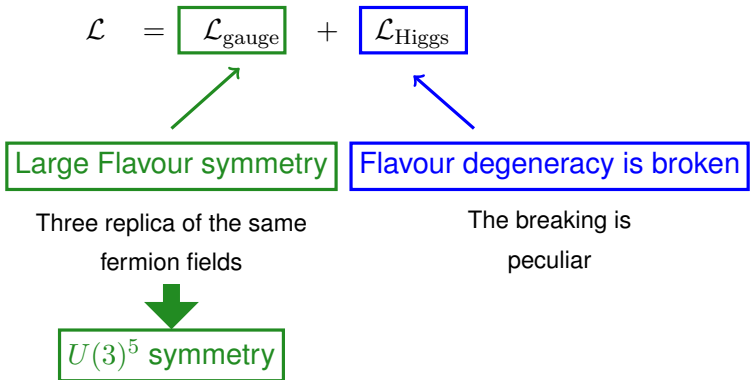
$$\mathcal{L}_{\text{Yukawa}} \supset Y_u^{ij} \bar{Q}_L^i H u_R^j$$

$$Y_u \sim y_t \begin{pmatrix} \text{light families} & \begin{pmatrix} 0.003 \\ 0.04 \end{pmatrix} \\ 1 \end{pmatrix}$$

$U(2)_u$ (red arrow pointing to the light families block)
 $U(2)_q$ (blue arrow pointing to the vector of values)

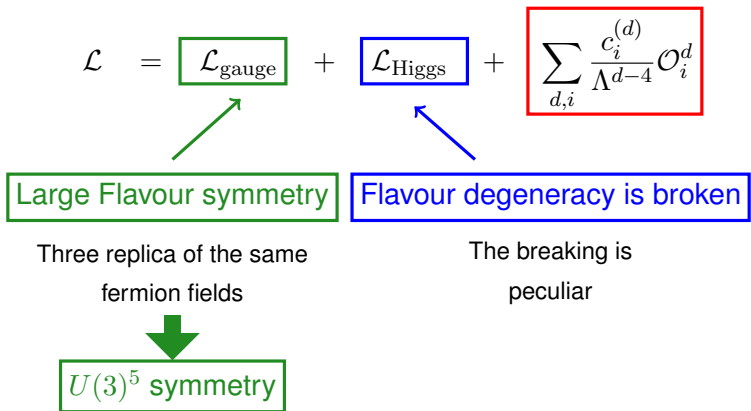
An approximate $U(2)^n$ is acting
on the light families!

The NP flavour problem



- In the SM: accidental $U(3)^5 \rightarrow \text{approx } U(2)^n$

The NP flavour problem

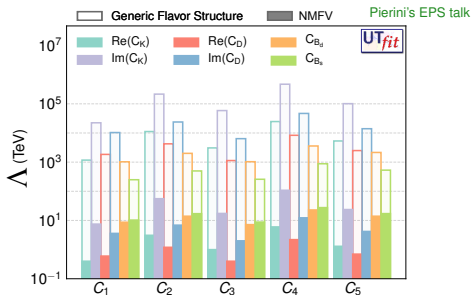
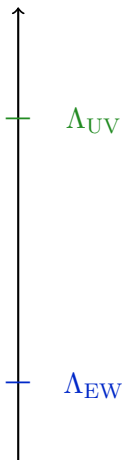


- In the SM: accidental $U(3)^5 \rightarrow \text{approx } U(2)^n$
- **What happens when we switch on NP?**

The NP flavour problem

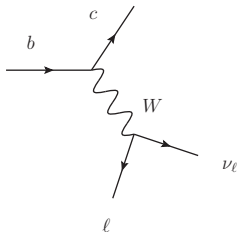
$$\mathcal{L} = \boxed{\mathcal{L}_{\text{gauge}}} + \boxed{\mathcal{L}_{\text{Higgs}}} + \boxed{\sum_{d,i} \frac{c_i^{(d)}}{\Lambda^{d-4}} \mathcal{O}_i^d}$$

- What is the energy scale of NP?
- Why haven't observed any violation of accidental symmetries yet?

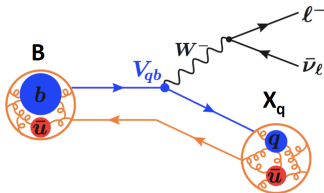


no breaking of the $U(2)^n$ flavour symmetry at low energies

Partonic vs Hadronic



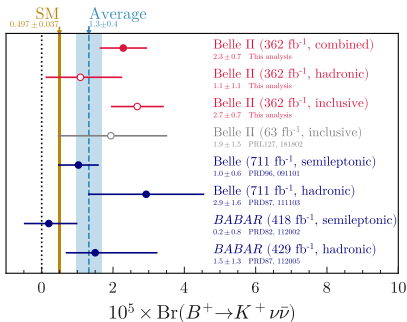
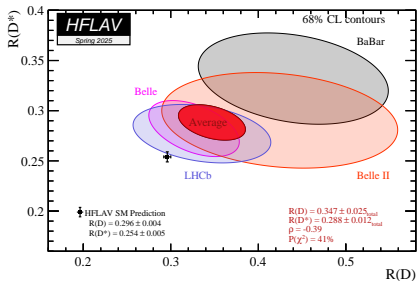
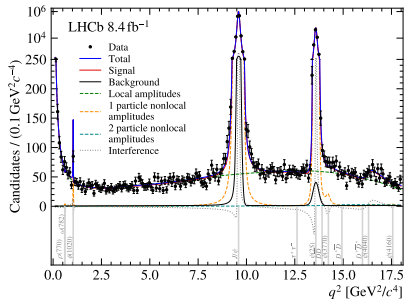
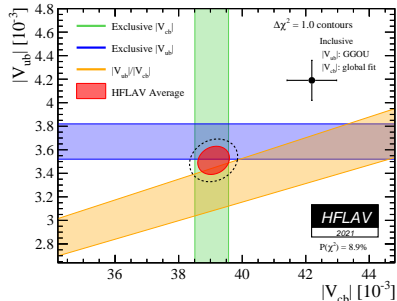
$$\mu_{\text{partonic}} = m_b$$



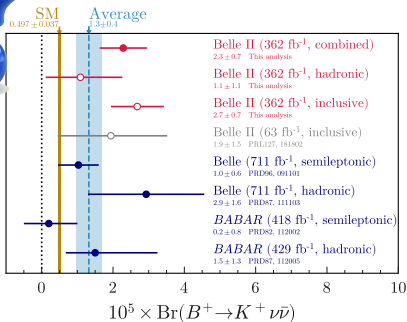
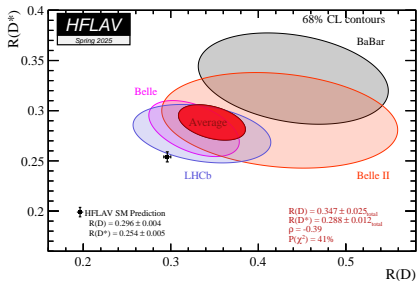
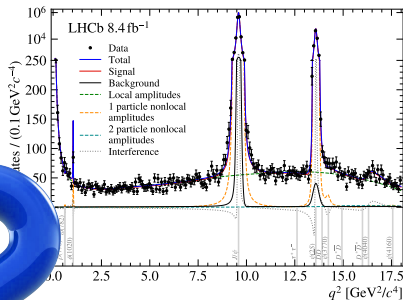
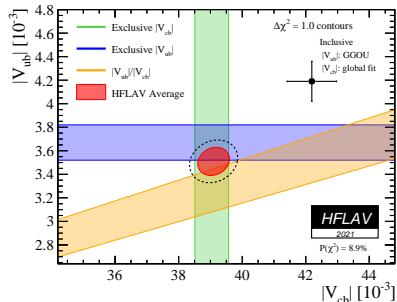
$$\mu_{\text{hadronic}} = \Lambda_{\text{QCD}}$$

**Fundamental challenge to match
partonic and hadronic descriptions**

Old and new puzzles in flavour physics



Old and new puzzles in flavour physics



The V_{cb} puzzle

The CKM matrix

Interaction basis

⇒ gauge interactions are diagonal

⇒ mass terms are not diagonal

$$-\mathcal{L}_Y = Y_d^{ij} \bar{Q}_L^i H d_R^j + Y_u^{ij} \bar{Q}_L^i \tilde{H} u_R^j + \text{h.c.}$$

Non-diagonal Yukawa

Mass basis

⇒ Yukawa couplings are diagonal

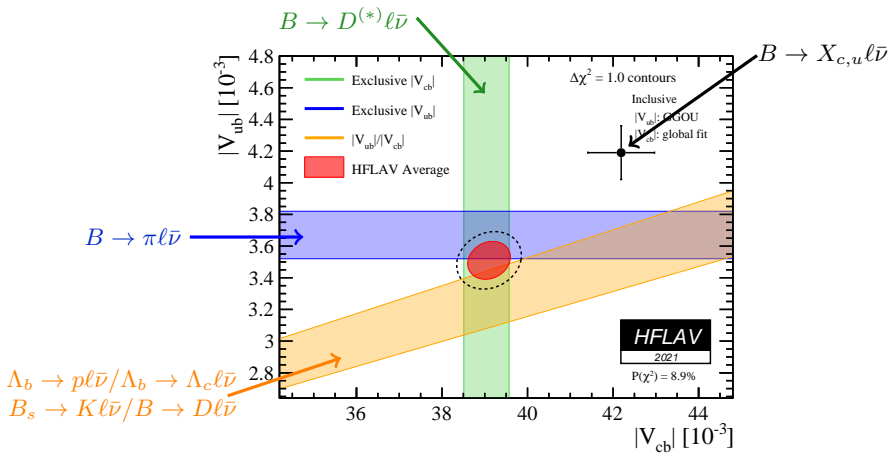
⇒ The CKM matrix is the remnant of the diagonalisation

$$\mathcal{L}_{cc} \propto \bar{u}_L^i \gamma^\mu d_L^j W_\mu^+ V_{ij}$$

CKM matrix

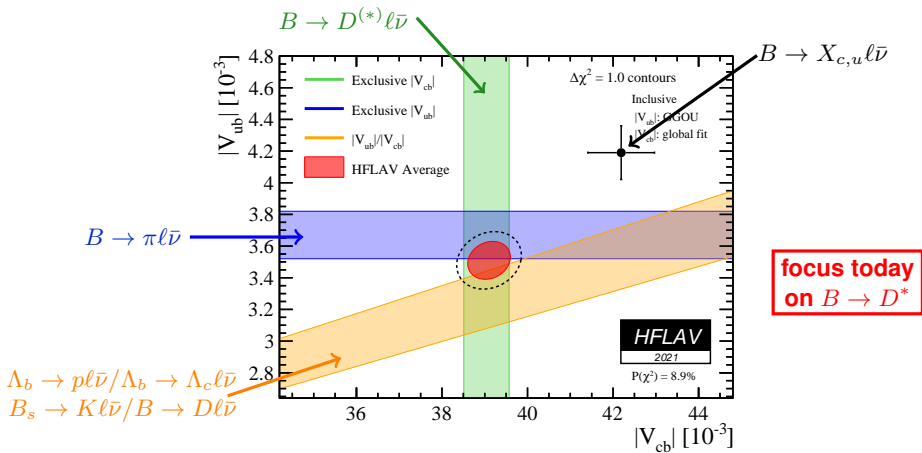
The $V_{cb} - V_{ub}$ puzzle

- Large discrepancies between inclusive and exclusive determinations
- Recent work mostly on $B \rightarrow D^*$ due to new lattice QCD form factors determinations
- When precision increases, more puzzles arise



The $V_{cb} - V_{ub}$ puzzle

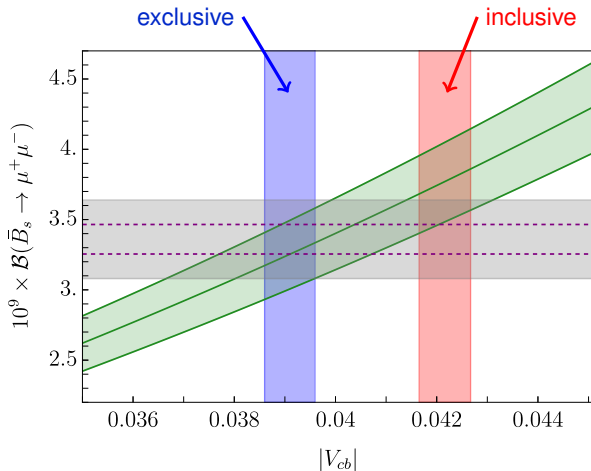
- Large discrepancies between inclusive and exclusive determinations
- Recent work mostly on $B \rightarrow D^*$ due to new lattice QCD form factors determinations
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Why is V_{cb} important?

$|V_{cb}|$ is a fundamental parameter to predict all flavour changing processes

$$\mathcal{B}(B_s \rightarrow \mu^+ \mu^-) \sim |V_{cb}|^2$$



Exclusive matrix elements

$$\langle H_c | J_\mu | H_b \rangle = \sum_i S_\mu^i \mathcal{F}_i$$

Exclusive matrix elements

$$\langle H_c | J_\mu | H_b \rangle = \sum_i S_\mu^i \mathcal{F}_i$$

Diagram illustrating the decomposition of the exclusive matrix element $\langle H_c | J_\mu | H_b \rangle$ into a sum over Lorentz structures S_μ^i and form factors \mathcal{F}_i .

Annotations:

- Red arrows pointing from Λ_{QCD} to S_μ^i and \mathcal{F}_i : scale Λ_{QCD}
- Green arrow pointing from S_μ^i to the text: independent Lorentz structures
- Blue arrow pointing from \mathcal{F}_i to the text: form factor

Exclusive matrix elements

$$\langle H_c | J_\mu | H_b \rangle = \sum_i S_\mu^i \mathcal{F}_i$$

Diagram illustrating the decomposition of the exclusive matrix element:

- \mathcal{F}_i is labeled "form factor" (indicated by a blue arrow).
- S_μ^i is labeled "independent Lorentz structures" (indicated by a green arrow).
- Λ_{QCD} is labeled "scale" (indicated by a red arrow pointing to the matrix element).

Form factors determinations

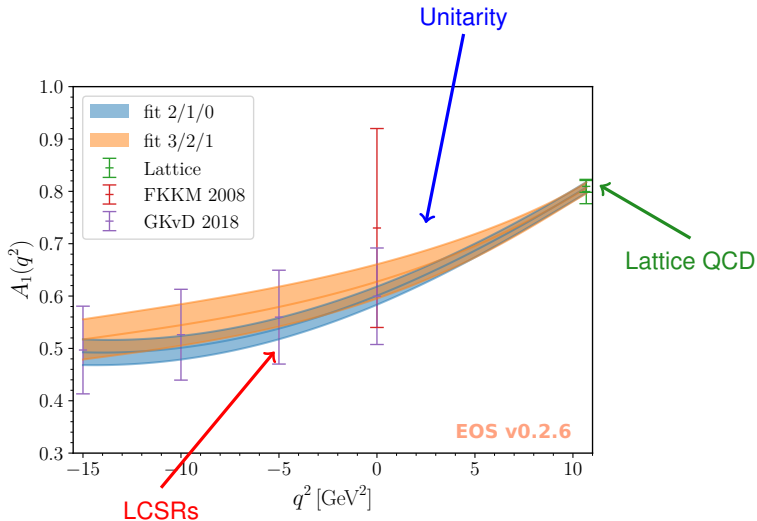
- Lattice QCD
- QCD SR, LCSR

only points at specific kinematic points

Form factors parametrisations

- HQET (CLN + improvements) \Rightarrow reduce independent degrees of freedom
- Analytic properties \rightarrow BGL

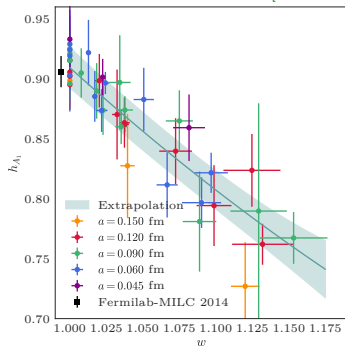
data points needed to fix the coefficients of the expansion



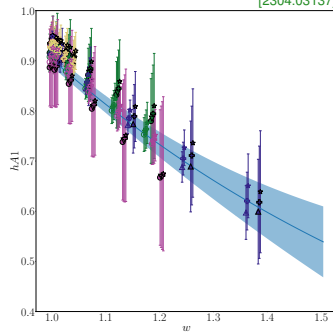
Other references: F. Bernlochner, Z. Ligeti, M. Papucci, M. Prim, D. Robinson, '22
P. Gambino, M. Jung, S. Schacht, '19

$B \rightarrow D^*$ from lattice away from zero recoil

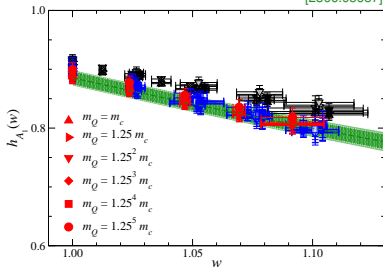
[2105.14019]



[2304.03137]

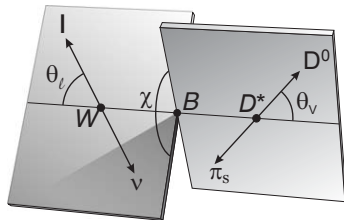
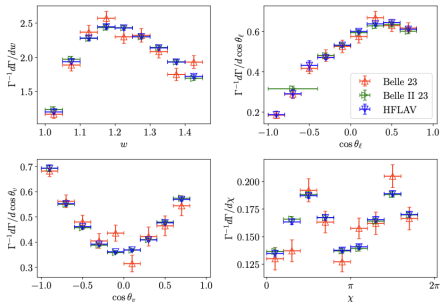


[2306.05657]



- Are these results compatible with each other?
- Are they compatible with experimental data?

New $B \rightarrow D^* \ell \bar{\nu}$ Belle and Belle II data



$$\frac{d\Gamma}{dw d\cos(\theta_\ell) d\cos(\theta_v) d\chi} = \frac{3G_F^2}{1024\pi^4} |V_{cb}|^2 \eta_{EW}^2 M_{B^*}^2 \sqrt{w^2 - 1} q^2$$

$$\times \left\{ (1 - \cos(\theta_\ell))^2 \sin^2(\theta_v) H_+^2(w) + (1 + \cos(\theta_\ell))^2 \sin^2(\theta_v) H_-^2(w) \right.$$

$$+ 4 \sin^2(\theta_\ell) \cos^2(\theta_v) H_0^2(w) - 2 \sin^2(\theta_\ell) \sin^2(\theta_v) \cos(2\chi) H_+(w) H_-(w)$$

$$- 4 \sin(\theta_\ell) (1 - \cos(\theta_\ell)) \sin(\theta_v) \cos(\theta_v) \cos(\chi) H_+(w) H_0(w)$$

$$+ 4 \sin(\theta_\ell) (1 + \cos(\theta_\ell)) \sin(\theta_v) \cos(\theta_v) \cos(\chi) H_-(w) H_0(w) \left. \right\}$$

- Between 7 to 10 bins per kinematic variable
- Available on HEPData with correlations
- Angular observables analysis are available, data just newly released

Analysis strategies

Setup

- BGL parametrisation
- Bayesian inference to apply unitarity

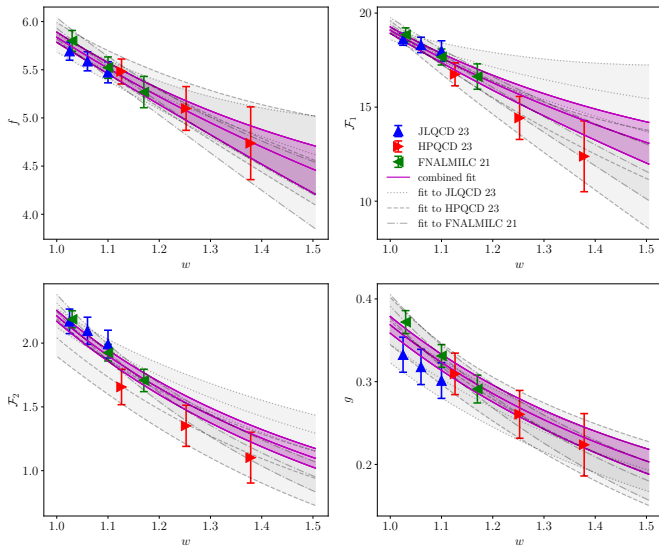
Flynn, Jüttner, Tsang, '23

Questions

- Combine the three LQCD datasets
 - ⇒ Is the combination acceptable?
- Combine with experimental data
- What are the consequences for phenomenology?

Lattice only

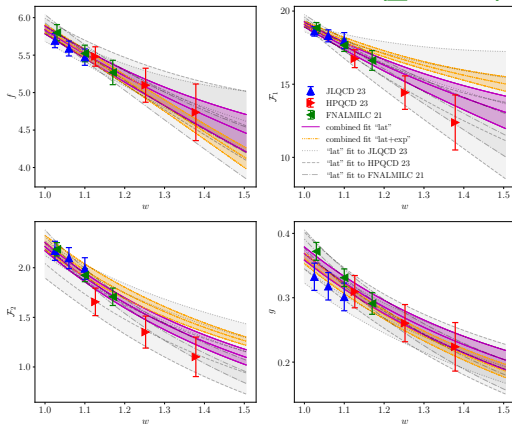
[MB, A. Jüttner, '24]



see also G. Martinelli, S. Simula, L. Vittorio, '23;24

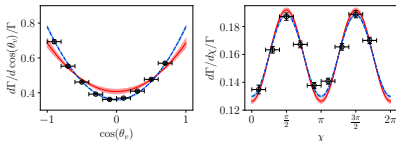
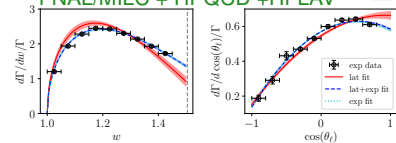
Lattice + experimental data

[MB, A. Jüttner, '24]

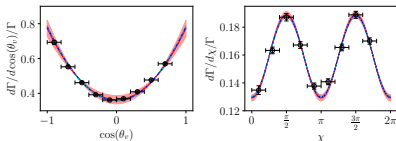
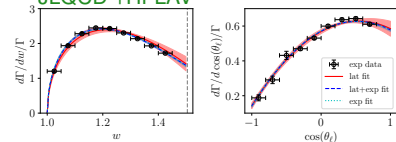


- Good fit quality (p -value $\sim 18\%$)
- Adding experimental data reduces the uncertainties, especially at large w
- Especially for F_1 and F_2 , the shape changes when including experimental data

FNAL/MILC + HPQCD + HFLAV [MB, A. Jüttner, '24]



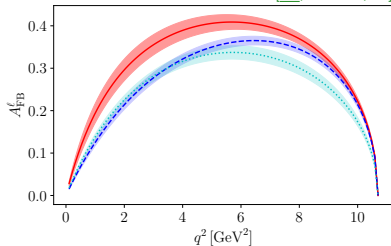
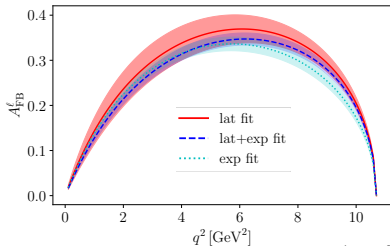
JLQCD + HFLAV



- Fit to HPQCD and FNAL/MILC misses experimental points
- BGL fit to experimental and lattice data has p -value $\sim 18\%$
- BGL coefficients shift of a few σ when including experimental data

Differential observables

[MB, A. Jüttner, '24]

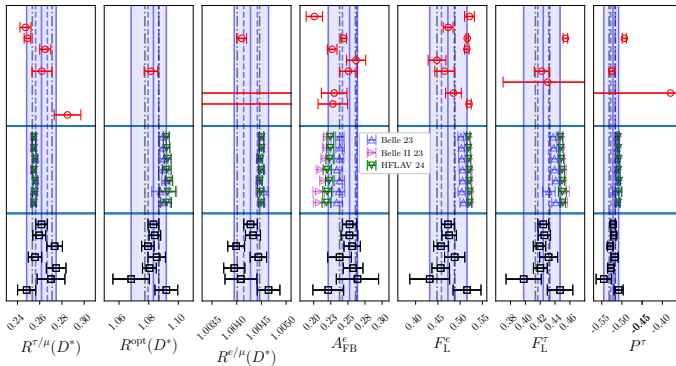


$$A_{\text{FB}}^{\ell} = \frac{\int_0^1 - \int_{-1}^0 d \cos \theta_{\ell} d\Gamma / d \cos \theta_{\ell}}{\int_0^1 + \int_{-1}^0 d \cos \theta_{\ell} d\Gamma / d \cos \theta_{\ell}}$$

- The combined lattice + experimental precision makes it possible to study the differences in the shape
- It is clear that there is a distinct difference between JLQCD and FNAL/MILC+HPQCD
- Difficult to understand what is going on, JLQCD errors are also a bit larger

Integrated observables

[MB, A. Jüttner, '24]



Bobeth et al. [71]
 Bordonone et al. [70]
 Bernlochner et al. [23]
 dispersive matrix Fedele et al. "lat+exp" [34]
 dispersive matrix Fedele et al. "lat" [34]
 dispersive matrix Martinelli et al. [33]
 LHCh [69]
 Belle [1, 67, 68]
 Belle II [3, 66]
 HFLAV Moriond 2024 (exp. average) [65]

exp+lat: JLQCD 23, HPQCD 23, FNALMILC 21
 exp+lat: JLQCD 23, FNALMILC 21
 exp+lat: FNALMILC 21, HPQCD 23
 exp+lat: JLQCD 23, HPQCD 23
 exp+lat: FNALMILC 21
 exp+lat: HPQCD 23
 exp+lat: JLQCD 23

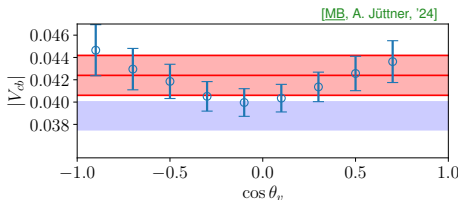
lat: JLQCD 23, HPQCD 23, FNALMILC 21
 lat: JLQCD 23, FNALMILC 21
 lat: FNALMILC 21, HPQCD 23
 lat: JLQCD 23, HPQCD 23
 lat: FNALMILC 21
 lat: HPQCD 23
 lat: JLQCD 23

- Significant scatter between various combinations of lattice results
 - We apply a systematic error to account for the spread
- Consistent scatter of the experimental results independently of the lattice information

⇒ see also: Fedele et al, '23

$|V_{cb}|$ extraction

$$|V_{cb}|_{\alpha,i} = \left(\Gamma_{\text{exp}} \left[\frac{1}{\Gamma} \frac{d\Gamma}{d\alpha} \right]_{\text{exp}}^{(i)} / \left[\frac{d\Gamma_0}{d\alpha}(\mathbf{a}) \right]_{\text{lat}}^{(i)} \right)^{1/2}$$



Blue band

- Frequentist fit p -value $\sim 0\%$
- Affected by d'Agostini Bias

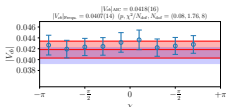
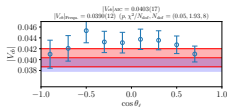
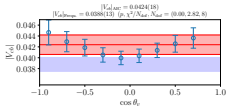
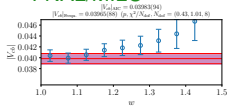
Red band

- Frequentist fit p -value $\sim 0\%$
- Akaike-Information-Criterion analysis: average over all possible fits with at least two data points and then weighted average

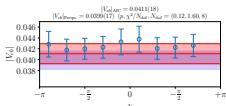
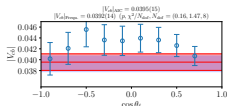
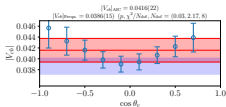
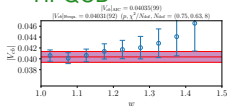
$$w_{\{\alpha,i\}} = \mathcal{N}^{-1} \exp \left(-\frac{1}{2} (\chi_{\{\alpha,i\}}^2 - 2N_{\text{dof},\{\alpha,i\}}) \right) \quad \text{where} \quad \mathcal{N} = \sum_{\text{sets } \{\alpha,i\}} w_{\text{set}}$$

$$|V_{cb}| = \langle |V_{cb}| \rangle \equiv \sum_{\text{sets } \{\alpha,i\}} w_{\text{set}} |V_{cb}|_{\text{set}}$$

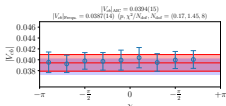
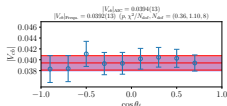
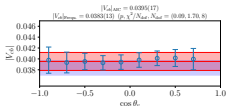
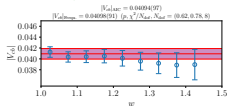
FNAL/MILC



HPQCD



JLQCD

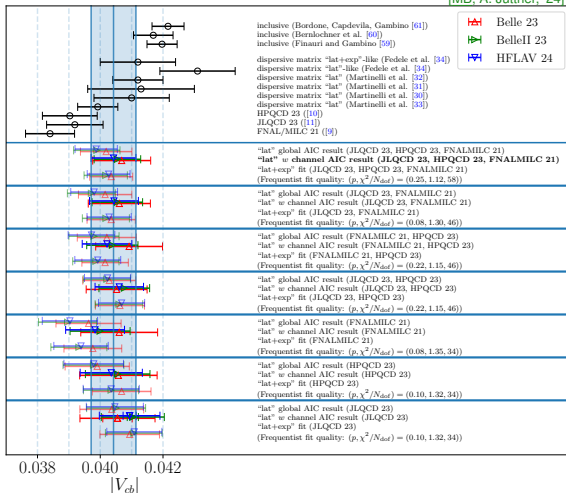


- The AIC nicely reduces the d'Agostini bias
- Some lattice data behave strangely
- Would it be safer to discard the angular distributions?
- Combining the three lattice datasets doesn't help, shape driven by FNAL/MILC and HPQCD

see also G. Martinelli, S. Simula, L. Vittorio, '23/24

$|V_{cb}|$ - Summary

[MB, A. Jüttner, '24]



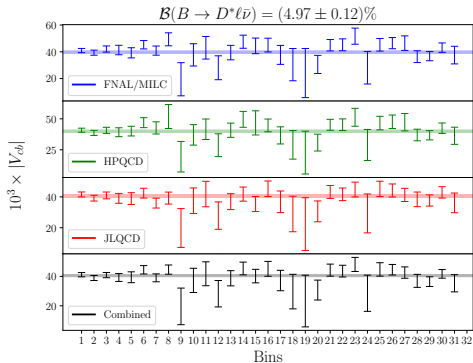
- Residual 2σ difference with inclusive
- The AIC produces slightly larger uncertainties, overall all results are quite consistent

What about other datasets?

[MB, O. Heald, A. Jüttner, in preparation]

Belle angular observables

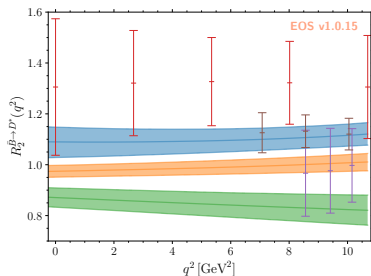
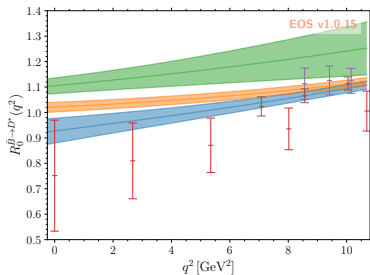
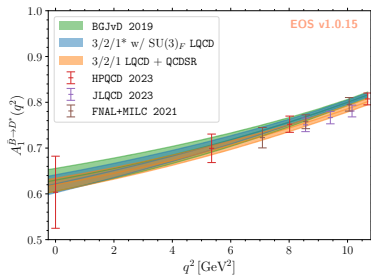
$$\frac{d\Gamma(\bar{B} \rightarrow D^* \ell \bar{\nu}_\ell)}{dw d\cos\theta_\ell d\cos\theta_V d\chi} = \frac{2G_F^2 \eta_{EW}^2 |V_{cb}|^2 m_B^4 m_{D^*}}{2\pi^4} \times \left(J_{1s} \sin^2 \theta_V + J_{1c} \cos^2 \theta_V \right. \\ + (J_{2s} \sin^2 \theta_V + J_{2c} \cos^2 \theta_V) \cos 2\theta_\ell + J_3 \sin^2 \theta_V \sin^2 \theta_\ell \cos 2\chi \\ + J_4 \sin 2\theta_V \sin 2\theta_\ell \cos \chi + J_5 \sin 2\theta_V \sin \theta_\ell \cos \chi + (J_{6s} \sin^2 \theta_V + J_{6c} \cos^2 \theta_V) \cos \theta_\ell \\ \left. + J_7 \sin 2\theta_V \sin \theta_\ell \sin \chi + J_8 \sin 2\theta_V \sin 2\theta_\ell \sin \chi + J_9 \sin^2 \theta_V \sin^2 \theta_\ell \sin 2\chi \right).$$



**Results are consistent with
our previous analysis**

Can a different parametrisation help?

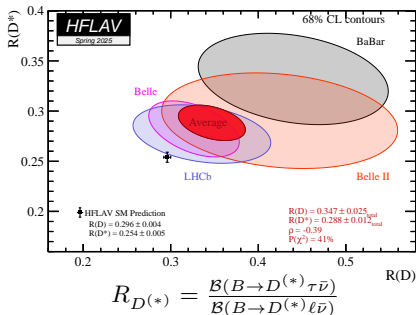
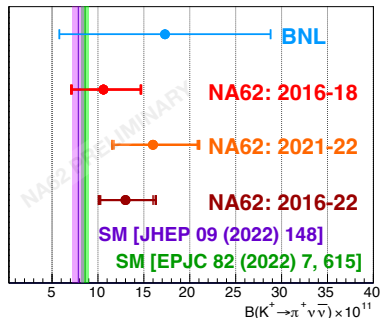
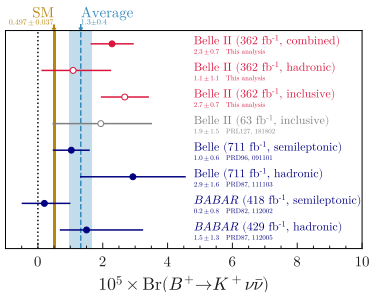
[MB, N. Gubernari, M. Jung D. van Dyk, '25]



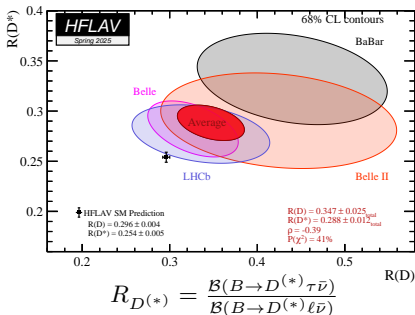
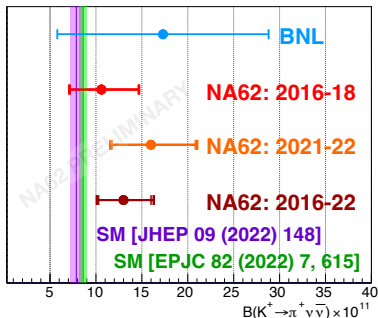
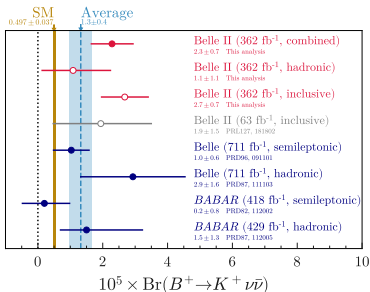
**same pattern
of deviations observed**

Outlook and prospects on BSM

What about BSM?



What about BSM?



Can we accommodate all these deviations together?

The EFT approach

- Since we haven't observed any clear sign of NP yet at low energies, we can work in an EFT context
 - ⇒ Agnostic of the nature of new physics, describe more than one UV model with the same operators
 - ⇒ Try to derive model-independent bounds
- We use the SMEFT
 - ⇒ Build all possible operators with SM fields and respecting SM symmetries
- The remnant of high-energy new physics is contained in the Wilson Coefficients
 - ⇒ With flavour, we have a lot of free degrees of freedom
 - ⇒ We need a criterium to infer their magnitude

The $U(2)^n$ symmetry for BSM

$$q_{3L} \sim (\mathbf{1}, \mathbf{1})$$

$$\ell_{3L} \sim (\mathbf{1}, \mathbf{1})$$

$$Q_L = (Q_L^1, Q_L^2) \sim (\bar{\mathbf{2}}, \mathbf{1})$$

$$L_L = (\ell_L^1, \ell_L^2) \sim (\mathbf{1}, \bar{\mathbf{2}})$$

Unbroken $U(2)^5$

$$Y_u = y_t \begin{pmatrix} 0 & 0 \\ 0 & 1 \end{pmatrix}$$



$$\bar{q}_{3L} \Gamma q_{3L} \quad \checkmark$$

$$\bar{q}_{3L} \Gamma Q \quad \times$$

The $U(2)^n$ symmetry for BSM

$$\begin{aligned}q_{3L} &\sim (\mathbf{1}, \mathbf{1}) \\ Q_L = (Q_L^1, Q_L^2) &\sim (\bar{\mathbf{2}}, \mathbf{1}) \\ V_q &\sim (\mathbf{2}, \mathbf{1})\end{aligned}$$

$$\begin{aligned}\ell_{3L} &\sim (\mathbf{1}, \mathbf{1}) \\ L_L = (\ell_L^1, \ell_L^2) &\sim (\mathbf{1}, \bar{\mathbf{2}}) \\ V_\ell &\sim (\mathbf{1}, \mathbf{2})\end{aligned}$$

Unbroken $U(2)^5$

$$Y_u = y_t \begin{pmatrix} 0 & 0 \\ 0 & 1 \end{pmatrix}$$



$$\bar{q}_{3L} \Gamma q_{3L} \quad \checkmark$$

$$\bar{q}_{3L} \Gamma Q \quad \times$$

Soft symmetry breaking

$$Y_u = y_t \begin{pmatrix} \Delta & V_q \\ 0 & 1 \end{pmatrix}$$



$$\bar{q}_{3L} \Gamma q_{3L} \quad \checkmark$$

$$\bar{q}_{3L} \Gamma (V_q Q) \quad \checkmark$$

Flavour Non-Universal New Physics

Dvali, Shifman, '00

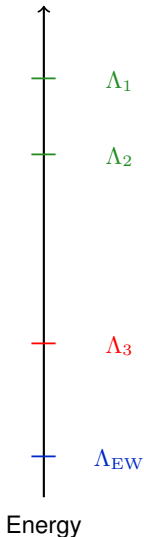
Panico, Pomarol, '16

MB, Cornella, Fuentes-Martin, Isidori '17

Allwicher, Isidori, Thomsen '20

Barbieri, Cornella, Isidori, '21

Davighi, Isidori '21



Basic idea:

- 1st and 2nd have small masses and small couplings to NP because they are generated by dynamics at a heavier scale
- 3rd generation is linked to dynamics at lower scales and has stronger couplings

Flavour deconstruction:

fermion families interact with different gauge groups and flavour hierarchies emerge as accidental symmetries

Flavour Non-Universal New Physics

Dvali, Shifman, '00

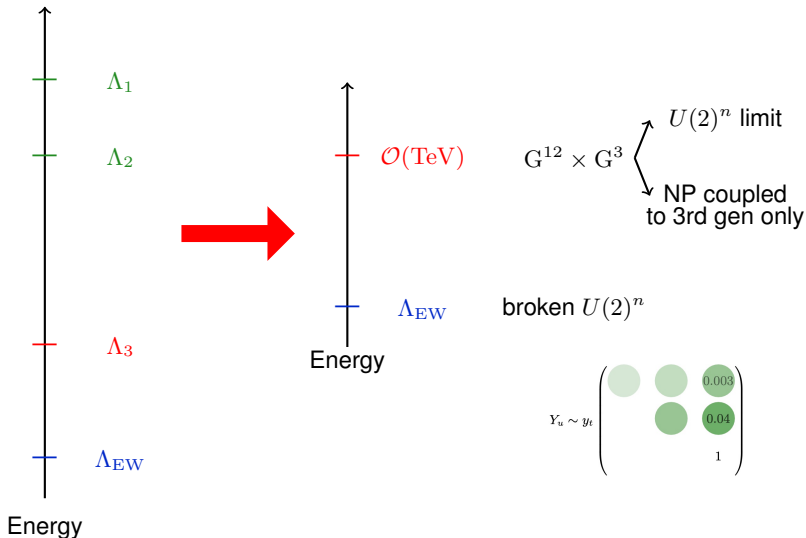
Panico, Pomarol, '16

MB, Cornella, Fuentes-Martin, Isidori '17

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Barbieri, Cornella, Isidori, '21

Davighi, Isidori '21




Which operators?

$$Q_{\ell q}^{\pm} = (\bar{q}_L^3 \gamma^{\mu} q_L^3)(\bar{\ell}_L^3 \gamma_{\mu} \ell_L^3) \pm (\bar{q}_L^3 \gamma^{\mu} \sigma^a q_L^3)(\bar{\ell}_L^3 \gamma_{\mu} \sigma^a \ell_L^3) \quad Q_S = (\bar{\ell}_L^3 \tau_R)(\bar{b}_R q_L^3)$$

Which operators?

$$Q_{\ell q}^{\pm} = (\bar{q}_L^3 \gamma^{\mu} q_L^3)(\bar{\ell}_L^3 \gamma_{\mu} \ell_L^3) \pm (\bar{q}_L^3 \gamma^{\mu} \sigma^a q_L^3)(\bar{\ell}_L^3 \gamma_{\mu} \sigma^a \ell_L^3) \quad Q_S = (\bar{\ell}_L^3 \tau_R)(\bar{b}_R q_L^3)$$


 $SU(2)$ singlet $SU(2)$ triplet scalar

Which operators?

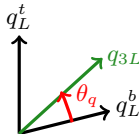
$$Q_{\ell q}^{\pm} = (\bar{q}_L^3 \gamma^{\mu} q_L^3)(\bar{\ell}_L^3 \gamma_{\mu} \ell_L^3) \pm (\bar{q}_L^3 \gamma^{\mu} \sigma^a q_L^3)(\bar{\ell}_L^3 \gamma_{\mu} \sigma^a \ell_L^3) \quad Q_S = (\bar{\ell}_L^3 \tau_R)(\bar{b}_R q_L^3)$$

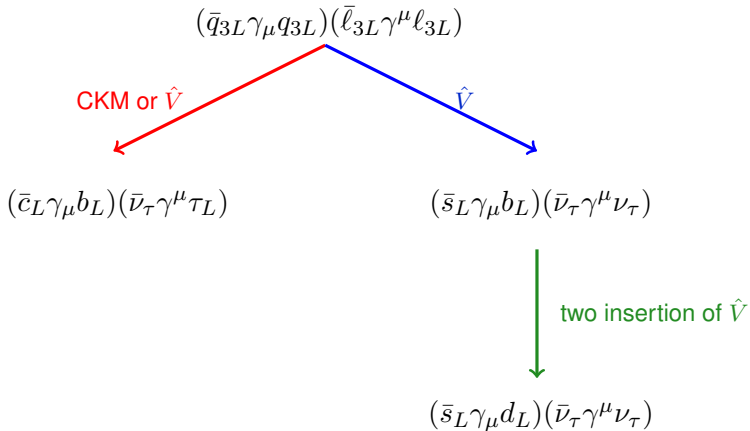
\uparrow
SU(2) singlet
 \uparrow
SU(2) triplet
 \uparrow
scalar

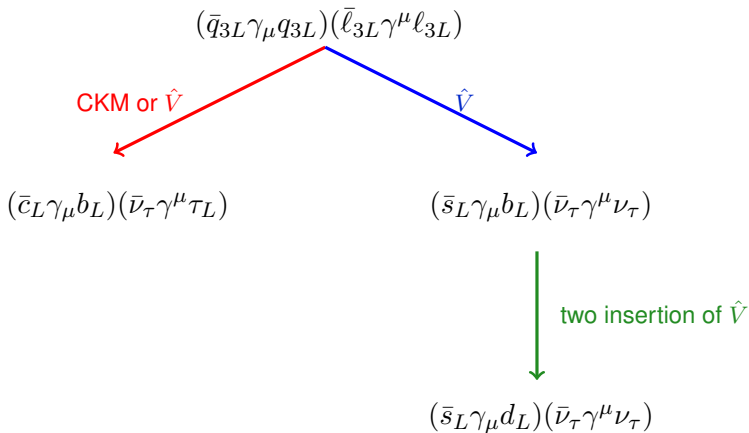
- Only left-handed neutrinos

- $q_{3L} \equiv q_L^b + \hat{V} \cdot Q_L$

$$q_L^b = \begin{pmatrix} V_{j3}^* u_L^j \\ b_L \end{pmatrix} \quad Q_L^i = \begin{pmatrix} V_{ji}^* u_L^j \\ d_L^i \end{pmatrix} \quad \hat{V}_q \equiv -\epsilon V_{ts} \begin{pmatrix} \kappa V_{td}/V_{ts} \\ 1 \end{pmatrix}$$







**Correlations among all these modes
is essential to prove NP scenarios**

What do we expect in the SMEFT?

$$\mathcal{L}_{\text{EFT}} \supset \frac{C_{bc\tau\tau}}{\Lambda^2} (\bar{b}_L \gamma_\nu c_L) (\bar{\nu}_\tau \gamma^\mu \tau_L)$$

From $U(2)^n \Rightarrow C_{bc\tau\tau} \sim V_{cb} \mathcal{O}(1)$

From $R_{D^{(*)}} \Rightarrow \Lambda \sim \mathcal{O}(\text{TeV})$

Using $SU(2)_L$ invariance, we have

$$\mathcal{L}_{\text{EFT}} \supset \frac{C_{ij\tau\tau}}{\Lambda^2} (\bar{d}_L^i \gamma_\nu d_L^j) (\bar{\nu}_\tau \gamma^\mu \nu_\tau)$$

$B^+ \rightarrow K^+ \nu \bar{\nu}$

From $U(2)^n \Rightarrow C_{bs\tau\tau} \sim V_{cb} \mathcal{O}(1)$

$K^+ \rightarrow \pi^+ \nu \bar{\nu}$

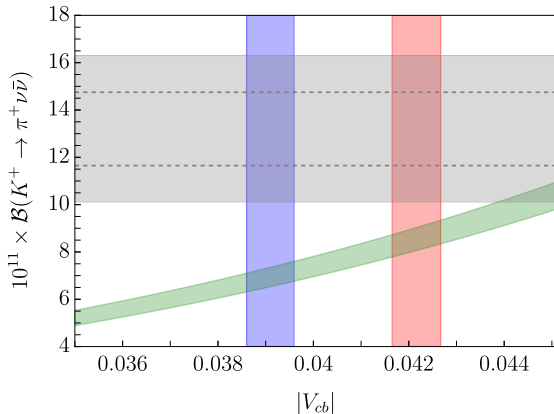
From $U(2)^n \Rightarrow C_{sd\tau\tau} \sim 10^{-1} V_{cb} \mathcal{O}(1)$

On the V_{cb} puzzle (again)

$$\mathcal{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu}) \propto |\lambda_{ts}|^2 \quad \lambda_{ts} \equiv \lambda |V_{cb}|^2 \left[(\bar{\rho} - 1) \left(1 - \frac{\lambda^2}{2} \right) + i\bar{\eta} \left(1 + \frac{\lambda^2}{2} \right) \right] + \mathcal{O}(\lambda^4)$$

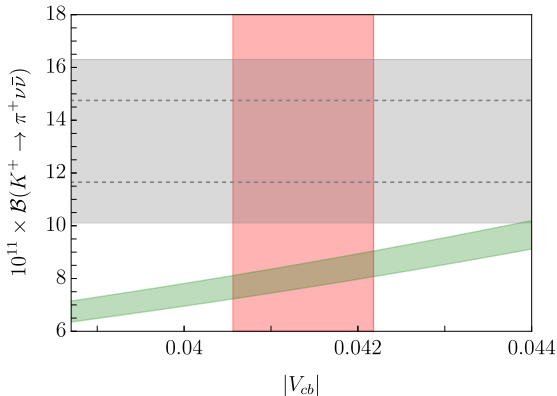
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$$\mathcal{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu}) \propto |\lambda_{ts}|^2 \quad \lambda_{ts} \equiv \lambda |V_{cb}|^2 \left[(\bar{\rho} - 1) \left(1 - \frac{\lambda^2}{2} \right) + i\bar{\eta} \left(1 + \frac{\lambda^2}{2} \right) \right] + \mathcal{O}(\lambda^4)$$



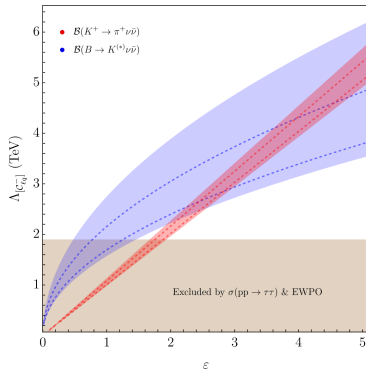
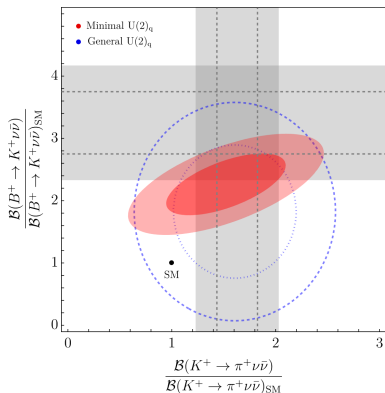
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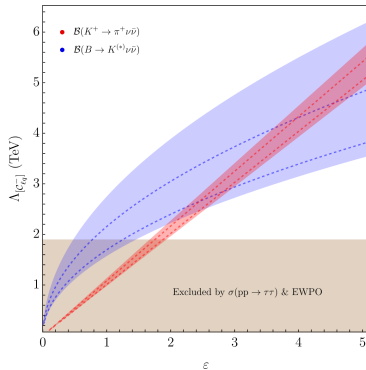
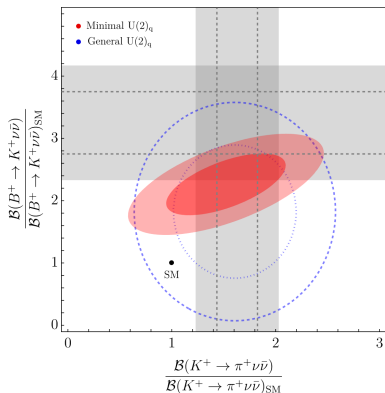


$$\mathcal{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu})^{\text{SM}} = (8.09 \pm 0.63) \times 10^{-11}$$

$$\mathcal{B}(K_L \rightarrow \pi^0 \nu \bar{\nu})^{\text{SM}} = (2.58 \pm 0.30) \times 10^{-11}$$



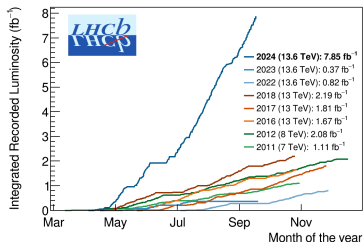
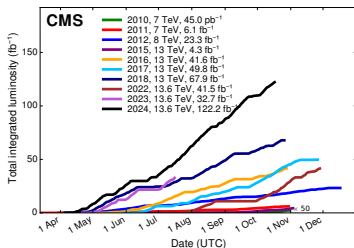
- The $U(2)^n$ symmetry creates a natural link between all this observables
- The complementarity between low- and high-energy data is useful to probe the parameter space



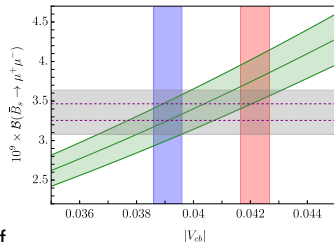
- The $U(2)^n$ symmetry creates a natural link between all this observables
- The complementarity between low- and high-energy data is useful to probe the parameter space

Further data is essential!

Experimental prospects



- Experimental facilities are delivering unprecedented datasets
- The experimental reach supported by new analysis techniques already superseded the expectations
- Theoretical advancements are crucial for achieving greater precision in understanding flavor processes and evaluating potential signs of new physics

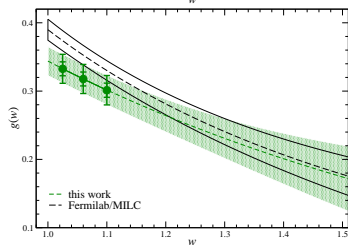
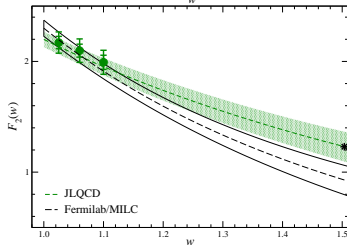
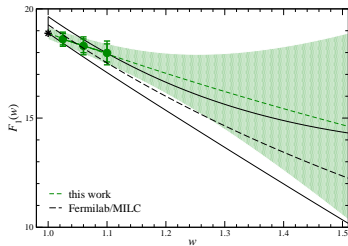
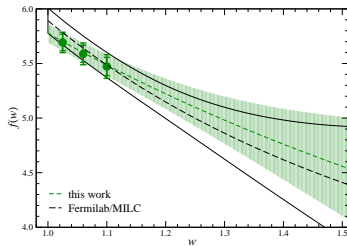


Summary

- Flavour physics has the potential to test for possible hints of extensions of the SM
- The main showstopper is the theoretical precision
- A lot of progress has been made, but a few pivotal puzzles persist
- There are hints for possible BSM directions, but more efforts and more data are needed to shed light on their nature

Appendix

Compatibility of lattice data



- Similar results with HPQCD
- There are some differences in the slopes

- How good is the compatibility?
- Do the differences yield significant pheno consequences?

Frequentist fit

K_f	$K_{\mathcal{F}_1}$	$K_{\mathcal{F}_2}$	K_g	$a_{g,0}$	$a_{g,1}$	$a_{g,2}$	$a_{g,3}$	p	χ^2/N_{dof}	N_{dof}
2	2	2	2	0.03138(87)	-0.059(24)	-	-	0.95	0.62	30
3	3	3	3	0.03131(87)	-0.046(36)	-1.2(1.8)	-	0.90	0.67	26
4	4	4	4	0.03126(87)	-0.017(48)	-3.7(3.3)	49.9(53.6)	0.79	0.75	22

- good fit quality
- lattice data are compatible
- no unitarity

Bayesian Fit

K_f	$K_{\mathcal{F}_1}$	$K_{\mathcal{F}_2}$	K_g	$a_{g,0}$	$a_{g,1}$	$a_{g,2}$	$a_{g,3}$
2	2	2	2	0.03018(76)	-0.101(21)	-	-
3	3	3	3	0.03034(78)	-0.087(24)	-0.34(45)	-
4	4	4	4	0.03035(77)	-0.089(23)	-0.27(41)	-0.04(45)

- unitarity regulates higher orders
- truncation dependent

What's the problem for BSM?

B-physics

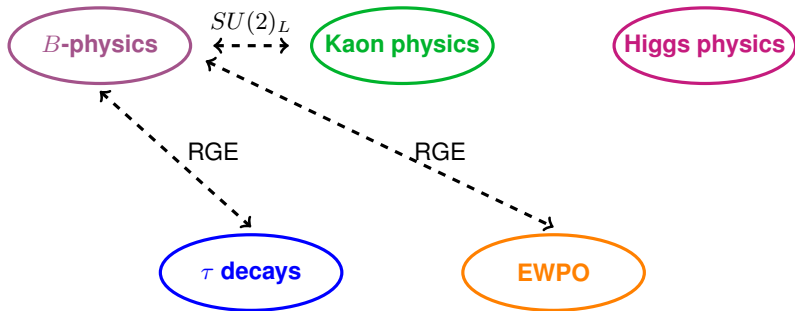
Kaon physics

Higgs physics

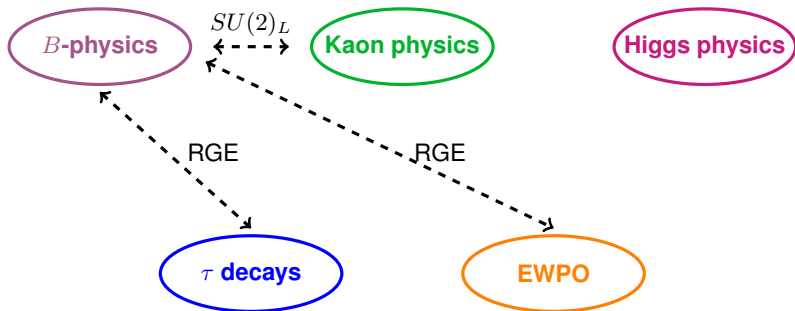
τ decays

EWPO

What's the problem for BSM?

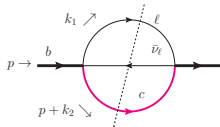


What's the problem for BSM?



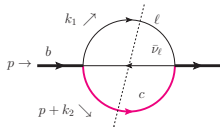
**How to satisfy all
the constraints
at the same time?**

Theory framework



$$\Gamma = \frac{1}{m_B} \text{Im} \int d^4x \langle B(p) | T \left\{ \mathcal{H}_{\text{eff}}^\dagger(x) \mathcal{H}_{\text{eff}}(0) \right\} | B(p) \rangle$$

Theory framework

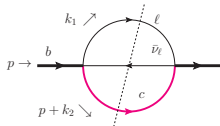


$$\Gamma = \frac{1}{m_B} \text{Im} \int d^4x \langle B(p) | T \left\{ \mathcal{H}_{\text{eff}}^\dagger(x) \mathcal{H}_{\text{eff}}(0) \right\} | B(p) \rangle$$

$$\sum_{n,i} \frac{1}{m_b^n} C_{n,i} \mathcal{O}_{n+3,i}$$

↑

Theory framework



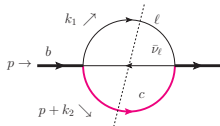
$$\Gamma = \frac{1}{m_B} \text{Im} \int d^4x \langle B(p) | T \left\{ \mathcal{H}_{\text{eff}}^\dagger(x) \mathcal{H}_{\text{eff}}(0) \right\} | B(p) \rangle$$

$$\sum_{n,i} \frac{1}{m_b^n} c_{n,i} \mathcal{O}_{n+3,i}$$

↑

- The Wilson coefficients are calculated perturbatively
- The matrix elements $\langle B(p) | \mathcal{O}_{n+3,i} | B(p) \rangle$ are non perturbative
 - ⇒ They need to be determined with non-perturbative methods, e.g. Lattice QCD
 - ⇒ They can be extracted from data
 - ⇒ With large n , large number of operators

Theory framework



$$\Gamma = \frac{1}{m_B} \text{Im} \int d^4x \langle B(p) | T \left\{ \mathcal{H}_{\text{eff}}^\dagger(x) \mathcal{H}_{\text{eff}}(0) \right\} | B(p) \rangle$$

$$\sum_{n,i} \frac{1}{m_b^n} c_{n,i} \mathcal{O}_{n+3,i}$$

- The Wilson coefficients are calculated perturbatively
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 - ⇒ They can be extracted from data
 - ⇒ With large n , large number of operators

↑
loss of predictivity

Theory framework for $B \rightarrow X_c \ell \bar{\nu}$

Double expansion in $1/m$ and α_s

$$\Gamma_{sl} = \Gamma_0 f(\rho) \left[1 + a_1 \left(\frac{\alpha_s}{\pi} \right) + a_2 \left(\frac{\alpha_s}{\pi} \right)^2 + a_3 \left(\frac{\alpha_s}{\pi} \right)^3 - \left(\frac{1}{2} - p_1 \left(\frac{\alpha_s}{\pi} \right) \right) \frac{\mu_\pi^2}{m_b^2} \right. \\ \left. + \left(g_0 + g_1 \left(\frac{\alpha_s}{\pi} \right) \right) \frac{\mu_G^2(m_b)}{m_b^2} + d_0 \frac{\rho_D^3}{m_b^3} - g_0 \frac{\rho_{LS}^3}{m_b^3} + \dots \right]$$

- The coefficients are known

- $\mu_\pi^2(\mu) = \frac{1}{2m_B} \langle B | \bar{b}_v (i\vec{D})^2 b_v | B \rangle_\mu \quad \mu_G^2(\mu) = \frac{1}{2m_B} \langle B | \bar{b}_v \frac{i}{2} \sigma_{\mu\nu} G^{\mu\nu} b_v | B \rangle_\mu$

⇒ No Lattice QCD determinations are available yet

- Use for the first time of α_s^3 corrections

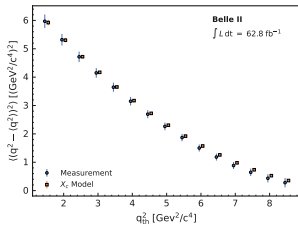
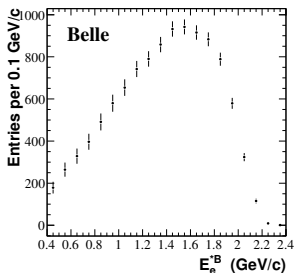
[Fael, Schönwald, Steinhauser, '20]

- Ellipses stands for higher orders

⇒ proliferation of terms and loss of predictivity

How do we constrain the hadronic parameters?

We need information from kinematic distributions

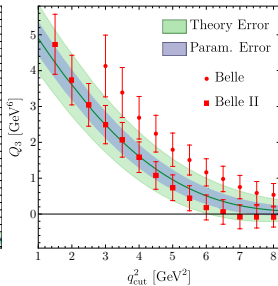
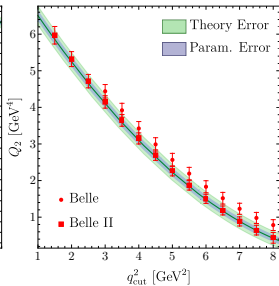
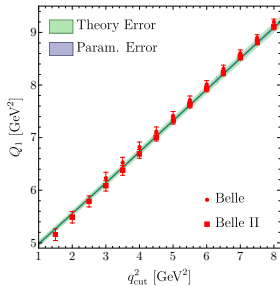


- Traditional method: Extract the hadronic parameters from moments of kinematic distributions in E_l and M_X
- New idea: Use q^2 moments to exploit the reduction of free parameters due to RPI
[Fael, Mannel, Vos, '18, Bernlochner et al, '22]
- Measurements of branching fractions are needed and are at the moment quite old
- Can we do it on the lattice?
[Gambino, Hashimoto, '20, '23, Hashimoto, Jüttner, et al, '23]

Global fit

[MB, Capdevila, Gambino, '21, Finauri, Gambino, '23]

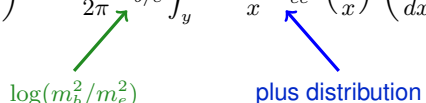
	m_b^{kin}	\overline{m}_c	μ_π^2	μ_G^2	ρ_D^3	ρ_{LS}^3	$10^2 \text{BR}_{c\ell\nu}$	$10^3 V_{cb} $	$\chi^2_{\text{min}}/(\text{dof})$
without	4.573	1.092	0.477	0.306	0.185	-0.130	10.66	42.16	22.3
q^2 -moments	0.012	0.008	0.056	0.050	0.031	0.092	0.15	0.51	0.474
Belle II	4.573	1.092	0.460	0.303	0.175	-0.118	10.65	42.08	26.4
	0.012	0.008	0.044	0.049	0.020	0.090	0.15	0.48	0.425
Belle	4.572	1.092	0.434	0.302	0.157	-0.100	10.64	41.96	28.1
	0.012	0.008	0.043	0.048	0.020	0.089	0.15	0.48	0.476
Belle &	4.572	1.092	0.449	0.301	0.167	-0.109	10.65	42.02	41.3
Belle II	0.012	0.008	0.042	0.048	0.018	0.089	0.15	0.48	0.559



Two calculation approaches

1. Splitting Functions

$$\left(\frac{d\Gamma}{dy}\right)^{(1)} = \frac{\alpha}{2\pi} \bar{L}_{b/e} \int_y^{1-\rho} \frac{dx}{x} P_{ee}^{(0)}\left(\frac{y}{x}\right) \left(\frac{d\Gamma}{dx}\right)^{(0)}$$



$\log(m_b^2/m_e^2)$ plus distribution

- Correction vanishes for the inclusive branching fraction
- Suitable for evaluating $\mathcal{O}(\alpha^2)$ and $\mathcal{O}(\alpha/m_b^n)$ corrections

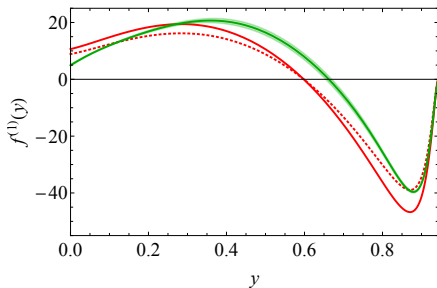
2. Full $\mathcal{O}(\alpha)$ corrections

- Access all corrections, not only the one that factorise
- Real corrections are computationally expensive
 - ⇒ Cuba library employed to carry out the 4-body integration
 - ⇒ Phase space splitting used to reduce the size of the integrands

Lepton Energy spectrum

[Bigi, MB, Gambino, Haisch, Piccione, '23]

- We compute bins in the lepton energy using the full $\mathcal{O}(\alpha)$ calculation
 - We compare them to the results given by the splitting functions
 - The difference the two calculations for the lepton energy spectrum and obtain a full analytic formula for the radiative corrections
- ⇒ Relatively small, easy-to-use formula to obtain branching fractions, lepton energy moments w/o cuts



$$f^{(1)}(y) = \frac{\bar{L}_{b/e}}{2} f_{LL}^{(1)}(y) + \Delta f^{(1)}(y)$$

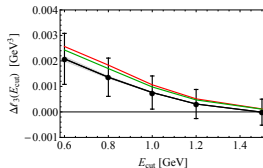
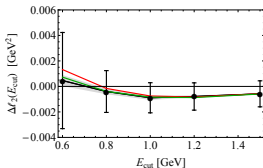
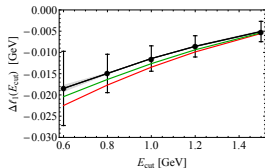
Comparison with data

[Bigi, MB, Gambino, Haisch, Piccione, '23]

- Babar provides data with and without applying PHOTOS to subtract QED effects

⇒ Perfect ground to test our calculations

⇒ Not the same for Belle at the moment, could be possible for future analysis

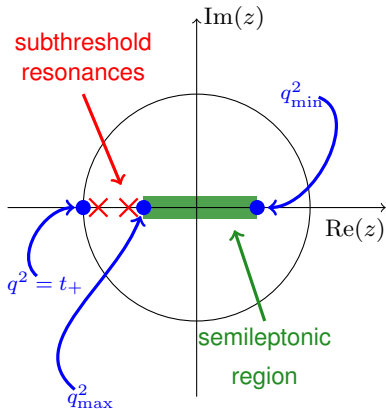


- The moments, since they are normalised, are not affected by the large threshold corrections
- The agreement with BaBar is very good

$$\langle E_\ell^n \rangle = \frac{\int_{E_\ell > E_{\ell, \text{cut}}} dE_\ell E_\ell^n \frac{d\Gamma}{dE_\ell}}{\Gamma_{E_\ell > E_{\ell, \text{cut}}}}$$

The z -expansion and unitarity

[Boyd, Grinstein, Lebed, '95, Caprini, Lellouch, Neubert, '98]



- in the complex plane form factors are real analytic functions
- q^2 is mapped onto the conformal complex variable z

$$z(q^2, t_0) = \frac{\sqrt{t_+ - q^2} - \sqrt{t_+ - t_0}}{\sqrt{t_+ - q^2} + \sqrt{t_+ - t_0}}$$

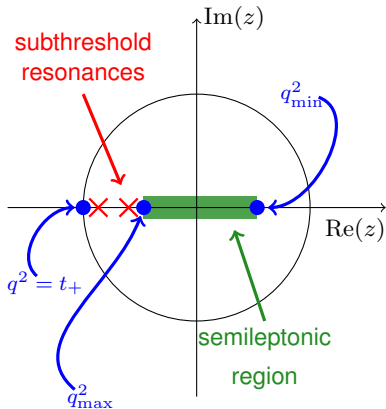
- q^2 is mapped onto a disk in the complex z plane, where $|z(q^2, t_0)| < 1$

$$F_i = \frac{1}{P_i(z)\phi_i(z)} \sum_{k=0}^{n_i} a_k^i z^k$$

$$\sum_{k=0}^{n_i} |a_k^i|^2 < 1$$

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BGL

How to apply unitarity

- Penalty function in the χ^2 or likelihood

[P. Gambino, M. Jung, S. Schacht, '19]

$$\chi^2 \rightarrow \chi^2(a_k^i, a_k^i|_{\text{data}}) + w_i \theta \left(\sum_{k=0}^{n_i} |a_k^i|^2 - 1 \right)$$

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- Dispersive Matrix Method

[M. Di Carlo, G. Martinelli, M. Naviglio, F. Sanfilippo, S. Simula, L. Vittorio, '21]

[G. Martinelli, S. Simula, L. Vittorio, '21,'23]

$$\mathbf{M} = \begin{pmatrix} \chi & \phi f & \phi_1 f_1 & \phi_2 f_2 & \dots & \phi_N f_N \\ \phi f & \frac{1}{1-z^2} & \frac{1}{1-zz_1} & \frac{1}{1-zz_2} & \dots & \frac{1}{1-zz_N} \\ \phi_1 f_1 & \frac{1}{1-z_1 z} & \frac{1}{1-z_1^2} & \frac{1}{1-z_1 z_2} & \dots & \frac{1}{1-z_1 z_N} \\ \phi_2 f_2 & \frac{1}{1-z_2 z} & \frac{1}{1-z_2 z_1} & \frac{1}{1-z_2^2} & \dots & \frac{1}{1-z_2 z_N} \\ \dots & \dots & \dots & \dots & \dots & \dots \\ \phi_N f_N & \frac{1}{1-z_N z} & \frac{1}{1-z_N z_1} & \frac{1}{1-z_N z_2} & \dots & \frac{1}{1-z_N^2} \end{pmatrix}$$

$$\det \mathbf{M} > 0 \Rightarrow \beta - \sqrt{\gamma} \leq f_0 \leq \beta + \sqrt{\gamma}$$

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- Bayesian inference

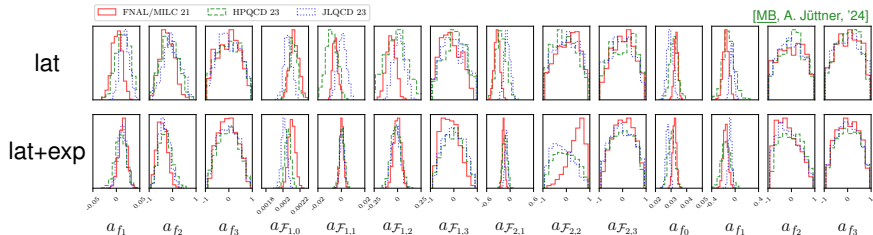
[J. Flynn, A. Jüttner, T. Tsang, '23]

$$\langle g(\mathbf{a}) \rangle = \mathcal{N} \int d\mathbf{a} g(\mathbf{a}) \pi(\mathbf{a}|\mathbf{f}, C_f) \pi_{\mathbf{a}}$$

$\theta(1 - |\mathbf{a}|^2)$

contains the lattice χ^2

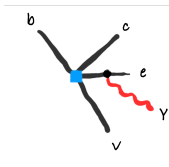
Posterior distribution



- Small shifts between lattice only and lattice + data
- Higher order coefficients well constrained by unitarity
- $a_{\mathcal{F}_{2,2}}$ has a strange behaviour, maybe kinematic constraints?

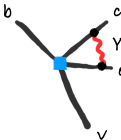
QED effects for inclusive V_{cb}

1. Collinear logs: captured by splitting functions



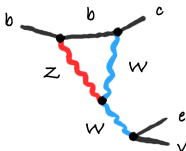
$$\sim \frac{\alpha_e}{\pi} \log^2 \left(\frac{m_b^2}{m_e^2} \right)$$

2. Threshold effects or Coulomb terms



$$\sim \frac{2\pi\alpha_e}{3}$$

3. Wilson Coefficient



$$\sim \frac{\alpha_e}{\pi} \left[\log \left(\frac{M_Z^2}{\mu^2} \right) - \frac{11}{6} \right]$$

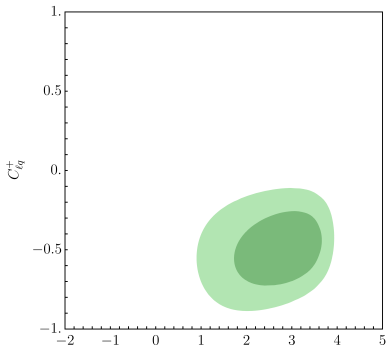
Branching ratio

[Bigi, MB, Gambino, Haisch, Piccione, '23]

- The total branching ratio is not affected by large logs due to KLN theorem
- The large corrections are from the Wilson Coefficient and the threshold effects

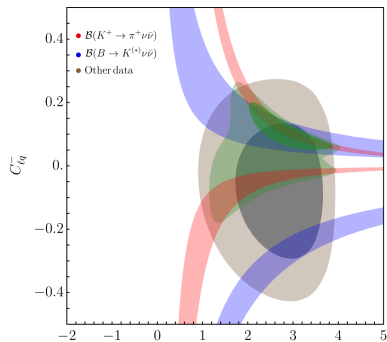
$$\frac{\Gamma}{\Gamma^{(0)}g(\rho)} = 1 + \frac{\alpha}{\pi} \left[\ln \left(\frac{M_Z^2}{m_b^2} \right) - \frac{11}{6} + 5.516(14) \right]$$
$$= 1 + \underbrace{(1.43\% - 0.44\%)}_{\text{Wilson Coefficient}} + \underbrace{1.32\%}_{\text{Threshold effects}} = 1 + 2.31\%$$

- Large shift of the branching ratio of the same order of the current error on V_{cb}
- How do we incorporate in the current datasets?
 - ⇒ Possible only on BaBar data
 - ⇒ A systematic approach is needed and foreseen for future experimental analysis
 - ⇒ How to evaluate structure-dependent terms is an open task



ϵ

- EWPO and direct searches
- $R_{D^{(*)}}$
- $B \rightarrow K^{(*)} \mu^+ \mu^-$



ϵ

- $B \rightarrow K^{(*)} \nu \bar{\nu}$
- $K^+ \rightarrow \pi^+ \nu \bar{\nu}$