

Ex 1: Partial Wave Expansion

1. $\vec{k} \cdot \vec{r} = kr \cos(\theta) = p \cos(\theta)$

Addition theorem for spherical harmonics:

$$\sum_{m=-\ell}^{\ell} \frac{4\pi}{2\ell+1} Y_{\ell m}^*(\theta_n, \phi_n) Y_{\ell m}(\theta_r, \phi_r) = P_{\ell}(\hat{k} \cdot \hat{r}) = P_{\ell}(\cos(\theta))$$

$$\Rightarrow 4\pi \sum_{m=-\ell}^{\ell} Y_{\ell m}^*(\theta_n, \phi_n) Y_{\ell m}(\theta_r, \phi_r) = (2\ell+1) P_{\ell}(\cos(\theta))$$

2.
$$\int_{-1}^1 d\cos(\theta) P_{\ell}(\cos(\theta)) \sum_{\ell=0}^{\infty} (2\ell+1) i^{\ell} j_{\ell}(p) P_{\ell}(\cos(\theta)) =$$

$$= \sum_{\ell=0}^{\infty} (2\ell+1) i^{\ell} j_{\ell}(p) \underbrace{\int_{-1}^1 d\cos(\theta) P_{\ell}(\cos(\theta)) P_{\ell}(\cos(\theta))}_{= \frac{2\delta_{\ell\ell}}{2\ell+1}} = 2i^{\ell} j_{\ell}(p)$$

For the other side of the equation, we have

$$\int_{-1}^1 d\cos(\theta) P_{\ell}(\cos(\theta)) e^{ip\cos(\theta)}$$

$$\xrightarrow{s=\cos(\theta)} j_{\ell}(p) = \frac{(-i)^{\ell}}{2} \int_{-1}^1 ds P_{\ell}(s) e^{ips}$$

3.
$$\frac{1}{2} \int_{-1}^1 ds e^{ips} = \frac{1}{2ip} (e^{ip} - e^{-ip}) = \frac{\sin(p)}{p}$$

To show eq. (7), we perform a proof by induction.

$\ell=1$:
$$(-p) \cdot \left(\frac{1}{p} \frac{d}{dp} \right) \frac{1}{2} \int_{-1}^1 ds e^{ips} =$$

$$= -\frac{i}{2} \int_{-1}^1 ds s e^{ips} = -\frac{i}{4} \left\{ \underbrace{\left[(s^2-1) e^{ips} \right]_{-1}^1}_{=0} - \int_{-1}^1 ds (s^2-1) i p e^{ips} \right\}$$

$$= \frac{(-p)^1}{2^{1+1} 1!} \int_{-1}^1 ds (s^2-1)^1 e^{ips}$$

$\ell \rightarrow \ell+1$:
$$j_{\ell+1}(p) = (-p)^{\ell+1} \left(\frac{1}{p} \frac{d}{dp} \right)^{\ell+1} \frac{\sin(p)}{p} =$$

$$= (-p)^{\ell+1} \left(\frac{1}{p} \frac{d}{dp} \right) \left(\frac{j_{\ell}(p)}{(-p)^{\ell}} \right) =$$

$$\stackrel{\text{induction hypothesis}}{=} \frac{(-p)^{\ell+1}}{2^{\ell+1} \ell!} \frac{1}{p} \frac{d}{dp} \int_{-1}^1 ds (s^2-1)^{\ell} e^{ips} =$$

$$= \frac{(-p)^{\ell+1}}{2^{\ell+1} \ell!} \frac{i}{p} \int_{-1}^1 ds (s^2-1)^{\ell} s e^{ips} =$$

$$= \frac{(-p)^{c+1}}{2^{c+1} e!} \cdot \frac{i}{p} \cdot \left\{ \underbrace{\left[\frac{(s^2-1)^{c+1}}{2(c+1)} e^{ips} \right]_{-1}^1}_{=0} - \int_{-1}^1 ds \frac{(s^2-1)^{c+1}}{2(c+1)} ip e^{ips} \right\} =$$

$$= \frac{(-p)^{c+1}}{2^{c+2}(c+1)!} \int_{-1}^1 ds (s^2-1)^{2c+1} e^{ips}$$

4. $j_c(p) = \frac{(-p)^c}{2^{c+1} e!} \int_{-1}^1 ds (s^2-1)^c e^{ips} =$

$$= \frac{i^c}{2^{c+1} e!} \int_{-1}^1 ds (s^2-1)^c \left(\frac{d^c}{ds^c} e^{ips} \right) =$$

$$= \frac{i^c}{2^{c+1} e!} \left\{ \left[(s^2-1)^c \frac{d^{c-1}}{ds^{c-1}} e^{ips} \right]_{-1}^1 - \int_{-1}^1 ds \left(\frac{d}{ds} (s^2-1)^c \right) \left(\frac{d^{c-1}}{ds^{c-1}} e^{ips} \right) \right\} =$$

$$= \frac{i^c}{2^{c+1} e!} (-1) \int_{-1}^1 ds \left(\frac{d}{ds} (s^2-1)^c \right) \left(\frac{d^{c-1}}{ds^{c-1}} e^{ips} \right) =$$

$$= \frac{i^c}{2^{c+1} e!} (-1)^c \int_{-1}^1 ds \left(\frac{d^c}{ds^c} (s^2-1)^c \right) e^{ips} =$$

repeat $c-1$ more times: the boundary term always takes the form

$$\left[\left(\frac{d^{n-1}}{ds^{n-1}} (s^2-1)^c \right) \cdot \left(\frac{d^{c-n}}{ds^{c-n}} e^{ips} \right) \right]_{-1}^1 = 0 \quad \forall n \leq c$$

$$= \frac{(-i)^c}{2} \int_{-1}^1 ds P_c(s) e^{ips}$$

Exercise 2: Phase Shifts for the potential well

1. For $r > R$, the Schrödinger equation for the particle is the one for a free particle with angular wavefunctions: $Y_{\ell m}(\theta, \phi)$ (spherical harmonics)
radial wavefunctions: $j_{\ell}(\rho)$, $y_{\ell}(\rho)$ (spherical Bessel functions)
 $\rho = k \cdot r$

So the most general form of the wave function for $r > R$ is

$$\Psi(r, \theta, \phi) = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \left(\underset{\text{same coefficients}}{a_{\ell m} j_{\ell}(k \cdot r) Y_{\ell m}(\theta, \phi) + b_{\ell m} y_{\ell}(k \cdot r) Y_{\ell m}(\theta, \phi)} \right)$$

However, we also know that the wave function of a free particle can be written as

$$\Psi(r, \theta, \phi) = A e^{i \vec{k} \cdot \vec{r}} + B e^{-i \vec{k} \cdot \vec{r}} = A e^{i k r \cos(\theta)} + B e^{-i k r \cos(\theta)}$$

which is independent of ϕ . This can only be true for the above expansion if

$$a_{\ell m}, b_{\ell m} \sim \delta_{m0}$$

We define $a_{\ell} \equiv a_{\ell 0}$, $b_{\ell} \equiv b_{\ell 0}$ and find

$$\Psi(r, \theta, \phi) = \sum_{\ell=0}^{\infty} (a_{\ell} j_{\ell}(k \cdot r) Y_{\ell 0}(\theta, \phi) + b_{\ell} y_{\ell}(k \cdot r) Y_{\ell 0}(\theta, \phi))$$

for $r > R$.

For $r < R$ the potential is a constant. Hence the solution to the Schrödinger equation is still the one for a free particle, but there are two differences

- as the Schrödinger equation is modified by the potential to

$$\begin{aligned} \nabla^2 \Psi(\vec{r}) &= - \underbrace{\frac{2m(E - V_0)}{\hbar^2}}_{= k'^2 - \frac{2mV_0}{\hbar^2} \equiv k'^2} \Psi(\vec{r}) \\ &= k'^2 - \frac{2mV_0}{\hbar^2} \equiv k'^2 \end{aligned}$$

$\Rightarrow$ we have to replace $k \rightarrow k'$ in the formula for the wave function

- the solution should be regular at $r=0$, but $y_{\ell}(x)$ diverges as $x \rightarrow 0$

$\Rightarrow$ all the coefficients for the y_{ℓ} vanish

To conclude, we have for $r < R$

$$\Psi(r, \theta, \phi) = \sum_{\ell=0}^{\infty} c_{\ell} j_{\ell}(k' \cdot r) Y_{\ell 0}(\theta, \phi)$$

2. As spherical harmonics are linearly independent functions, we can impose continuity and differentiability term by term at $r=R$, so for all ℓ , we have

$$c_\ell j_\ell(k'R) = a_\ell j_\ell(kR) + b_\ell y_\ell(kR) \quad (\text{cont.})$$

$$c_\ell k' j'_\ell(k'R) = a_\ell k j'_\ell(kR) + b_\ell k y'_\ell(kR) \quad (\text{diff.})$$

$$\Rightarrow \frac{a_\ell j_\ell(kR) + b_\ell y_\ell(kR)}{j_\ell(k'R)} = \frac{a_\ell k j'_\ell(kR) + b_\ell k y'_\ell(kR)}{k' j'_\ell(k'R)}$$

$$\Rightarrow \frac{b_\ell}{a_\ell} = \frac{k j'_\ell(kR) j_\ell(k'R) - k' j'_\ell(k'R) j_\ell(kR)}{k' j'_\ell(k'R) y_\ell(kR) - k y'_\ell(kR) j_\ell(k'R)}$$

$$3. \psi(r, \theta, \phi) \xrightarrow{r \rightarrow \infty} \sum_{\ell=0}^{\infty} \frac{1}{kr} \left[a_\ell \cos(kr - \frac{(\ell+1)\pi}{2}) + b_\ell \sin(kr - \frac{(\ell+1)\pi}{2}) \right] Y_{\ell m}(\theta, \phi)$$

$$= \sum_{\ell=0}^{\infty} \frac{1}{2kr} \left[(a_\ell - ib_\ell) e^{ikr - i(\ell+1)\frac{\pi}{2}} + (a_\ell + ib_\ell) e^{-ikr + i(\ell+1)\frac{\pi}{2}} \right] Y_{\ell m}(\theta, \phi)$$

$$\Rightarrow a_\ell - ib_\ell = e^{i\delta_\ell} \quad \text{and} \quad a_\ell + ib_\ell = e^{-i\delta_\ell} \quad (\text{consistent})$$

Dividing the two, we have

$$e^{2i\delta_\ell} = \frac{a_\ell - ib_\ell}{a_\ell + ib_\ell} = \frac{1 - i\alpha_\ell}{1 + i\alpha_\ell}$$

$$4. j_0(p) = \frac{\sin(p)}{p}, \quad j'_0(p) = \frac{p \cos(p) - \sin(p)}{p^2}$$

$$y_0(p) = -\frac{\cos(p)}{p}, \quad y'_0(p) = \frac{p \sin(p) + \cos(p)}{p^2}$$

$$\Rightarrow \alpha_0 = \frac{(kR \cos(kR) - \sin(kR)) \sin(k'R) - (k'R \cos(k'R) - \sin(k'R)) \sin(kR)}{(k'R \cos(k'R) - \sin(k'R)) (-\cos(kR)) - (kR \sin(kR) + \cos(kR)) \sin(k'R)}$$

$$= \frac{k' \cos(k'R) \sin(kR) - k \cos(kR) \sin(k'R)}{k' \cos(k'R) \cos(kR) + k \sin(kR) \sin(k'R)}$$

We expand for small potential $V_0 \ll \frac{\hbar^2 k^2}{2m}$:

$$k' \approx \sqrt{k^2 \left(1 - \frac{2mV_0}{\hbar^2 k^2}\right)} \approx k - \frac{mV_0}{\hbar^2 k} + \mathcal{O}(V_0^2)$$

$$\cos(k'R) \approx \cos(kR) + \frac{mV_0}{\hbar^2 k} R \sin(kR) + \mathcal{O}(V_0^2)$$

$$\sin(k'R) \approx \sin(kR) - \frac{mV_0}{\hbar^2 k} R \cos(kR) + \mathcal{O}(V_0^2)$$

$$\Rightarrow \alpha_0 \approx \frac{kR - \cos(kR) \sin(kR)}{\cos^2(kR) + \sin^2(kR)} \cdot \frac{mV_0}{\hbar^2 k^2} + \mathcal{O}(V_0^2)$$

$$= kR - \cos(kR) \sin(kR)$$

5. The scattering amplitude is given by

$$\begin{aligned}
 f_n(\theta, \phi) &= \frac{1}{k} \sum_{\ell=0}^{\infty} (2\ell+1) e^{i\delta_\ell} \sin(\delta_\ell) P_\ell(\cos(\theta)) = \\
 &= \frac{1}{k} \sum_{\ell=0}^{\infty} (2\ell+1) \frac{1}{2i} (e^{2i\delta_\ell} - 1) P_\ell(\cos(\theta)) = \\
 &= \frac{1}{k} \frac{1}{2i} (e^{2i\delta_0} - 1) + \frac{1}{k} \sum_{\ell=1}^{\infty} (2\ell+1) e^{i\delta_\ell} \sin(\delta_\ell) P_\ell(\cos(\theta)) \\
 \frac{1}{2i} (e^{2i\delta_0} - 1) &= \frac{1}{2i} \left(\frac{1-i\alpha_0}{1+i\alpha_0} - 1 \right) = -\frac{\alpha_0}{1-i\alpha_0} = \\
 &\stackrel{\uparrow}{=} -\alpha_0 + \sigma(V_0^2) = (kR - \cos(kR)\sin(kR)) \frac{mV_0}{\hbar^2 k^2} + O(V_0^2) \\
 \alpha_0 &= \sigma(V_0)
 \end{aligned}$$

So we have

$$\begin{aligned}
 f_n(\theta, \phi) &= -\frac{\alpha_0}{k} + (\text{partial waves } \ell \geq 1) = \\
 &= -\frac{mV_0}{\hbar^2} \frac{kR - \cos(kR)\sin(kR)}{k^3} + O(V_0^2) + (\text{p.w. } \ell \geq 1)
 \end{aligned}$$

6. The contribution of only the $\ell=0$ partial wave to the cross section is

$$\begin{aligned}
 \sigma_{\ell=0} &= \int d\Omega \left| -\frac{mV_0}{\hbar^2} \cdot \frac{kR - \cos(kR)\sin(kR)}{k^3} \right|^2 = \\
 &= 4\pi \frac{m^2 V_0^2}{\hbar^4} \cdot \frac{(kR - \cos(kR)\sin(kR))^2}{k^6} = \\
 &= \frac{\pi m^2 V_0^2}{\hbar^4} \frac{(2kR - \sin(2kR))^2}{k^6} = \\
 &= \frac{\pi m^2 V_0^2}{\hbar^4} \frac{\left(\frac{(2kR)^3}{6}\right)^2 + O((kR)^2)}{k^6} = \\
 &= \frac{16\pi m^2 V_0^4}{9\hbar^4} R^6 + O((kR)^2)
 \end{aligned}$$

This is exactly the same as the full cross section, so in the low energy limit, the contribution of the $\ell=0$ partial wave to the cross section dominates.

Exercise 3: Sommerfeld enhancement

1. Define $u_{\kappa, \ell}(r) \equiv r R_{\kappa, \ell}(r)$.

For a potential blowing up no faster than $\frac{1}{r^2}$ at the origin, the asymptotic behaviour is governed by the dominant angular term,

$$\frac{\hbar^2}{2m} \frac{d^2 u_{\kappa, \ell}(r)}{dr^2} = \frac{\hbar^2}{2m} \frac{\ell(\ell+1)}{r^2} u_{\kappa, \ell}(r), \quad \ell \neq 0$$

The general solution is

$$u_{\kappa, \ell}(r) = C_1 r^{\ell+1} + C_2 r^{-\ell}$$

But demanding regularity at $r=0$, we have $C_2=0$.

So $u_{\kappa, \ell}(r) \xrightarrow{r \rightarrow 0} r^{\ell+1}$, which implies $R_{\kappa, \ell}(r) \xrightarrow{r \rightarrow 0} r^{\ell}$.

Note: for $\ell=0$, the asymptotic behaviour depends on the potential!

2. We have

$$\begin{aligned} \psi_{\kappa}(\vec{r}) &= \frac{1}{\kappa} \sum_{\ell=0}^{\infty} (2\ell+1) i^{\ell} e^{i\delta_{\ell}} P_{\ell}(\cos(\theta)) R_{\kappa, \ell}(r) = \\ &= \frac{1}{\kappa} e^{i\delta_0} R_{\kappa, 0}(r) + \underbrace{\frac{1}{\kappa} \sum_{\ell=1}^{\infty} (2\ell+1) i^{\ell} e^{i\delta_{\ell}} P_{\ell}(\cos(\theta)) R_{\kappa, \ell}(r)}_{\text{goes to zero as } r \rightarrow 0} \\ &\quad \text{as } R_{\kappa, \ell}(r) \xrightarrow{r \rightarrow 0} r^{\ell} \end{aligned}$$

$$\Rightarrow \psi_{\kappa}(0) = \frac{1}{\kappa} e^{i\delta_0} R_{\kappa, 0}(0)$$

So the Sommerfeld enhancement factor reads

$$S_{\kappa} = |\psi_{\kappa}(0)|^2 = \left| \frac{R_{\kappa, 0}(0)}{\kappa} \right|^2.$$

3. For the Coulomb potential, we have

$$S_{\kappa} = \left| \frac{\eta}{e^{\eta} - 1} \right| = \frac{\eta}{e^{\eta} - 1}$$

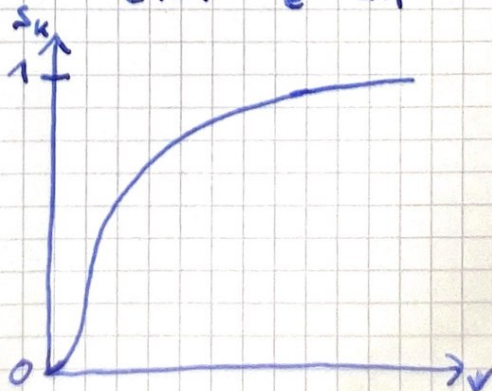
Case 1: repulsive potential $\alpha > 0$

$$\Rightarrow \eta = \frac{2\pi\alpha}{v} > 0 \quad \Rightarrow \quad S_{\kappa} = \frac{\eta}{e^{\eta} - 1} = \frac{\frac{2\pi\alpha}{v}}{e^{\frac{2\pi\alpha}{v}} - 1}$$

qualitative behaviour:

$$S_{\kappa} \xrightarrow{v \rightarrow 0} 0$$

$$S_{\kappa} \xrightarrow{v \rightarrow \infty} 1$$

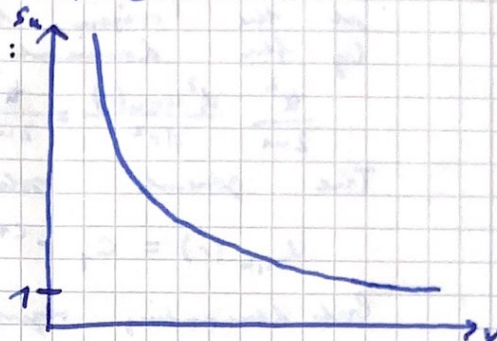


Case 2: attractive potential $\alpha < 0$
 $\Rightarrow \eta < 0 \Rightarrow S_u = \frac{-\frac{2\pi\alpha}{v}}{1 - e^{\frac{2\pi\alpha}{v}}}$

qualitative behavior:

$$S_u \xrightarrow{v \rightarrow 0} \infty$$

$$S_u \xrightarrow{v \rightarrow \infty} 1$$



Discussion:

- for a high velocity particle, $v \rightarrow c$, the impact of the potential becomes small and $S_u \sim 1$ in both the attractive and repulsive case (a proper treatment requires a relativistic framework, however)
- for a low velocity particle, the potential has large effects on the interaction probability
 - in the repulsive case: $S_u \rightarrow 0$ as $v \rightarrow 0$, i.e. the repulsion makes it unlikely for the particles to interact
 - in the attractive case: $S_u \rightarrow \infty$ as $v \rightarrow 0$, i.e. the attraction enhances the interaction probability drastically