

Density of states 1

To compute transition rates to final states of nearly degenerate energies $\rightarrow$ density of final states.

Consider, for example, the emission of a particle $\rightarrow$
 $\rightarrow$ free particle

$$\langle x | p \rangle = A e^{\frac{i}{\hbar} \vec{p} \cdot \vec{x}}$$

In a box of side $L \Rightarrow A = \frac{1}{L^{3/2}}$

And then we impose periodic boundary c.

$$\Psi(x+L, y, z) = \Psi(x, y, z), \quad y+L, \quad z+L, \dots \Rightarrow$$

$$e^{i p_x L / \hbar} = e^{i p_y L / \hbar} = e^{i p_z L / \hbar} = 1$$

which means

$$p_x = \frac{2\pi\hbar}{L} n_1, \quad p_y = \frac{2\pi\hbar}{L} n_2, \quad p_z = \frac{2\pi\hbar}{L} n_3$$

In a transition rate $\Gamma = \sum_{n_1} \sum_{n_2} \sum_{n_3} \Gamma_{i \rightarrow f(n_1, n_2, n_3)} \xrightarrow{\text{continuum limit}} \int d^3 n \Gamma$

$$d^3 n = dn_1 dn_2 dn_3 = \frac{V}{(2\pi\hbar)^3} dp_x dp_y dp_z$$

$$E = \frac{\vec{p}^2}{2m} \Rightarrow \frac{V}{(2\pi\hbar)^3} d^3 \vec{p} = \frac{V}{(2\pi\hbar)^3} d\Omega p^2 dp = \frac{V}{(2\pi\hbar)^3} p m dE d\Omega$$

$$d^3n = \frac{V}{(2\pi\hbar)^3} p \cdot m dE d\Omega \Rightarrow \rho(E) = \frac{d^3n}{dE} = \frac{V}{(2\pi\hbar)^3} p \cdot m d\Omega$$

(The density increases with V , correctly, more states per energy unit if V increases, see quantisation of p_s).

For a relativistic particle

$$\hbar\omega = E = pc \Rightarrow \frac{V}{(2\pi\hbar)^3} d^3\vec{p} = \frac{V}{(2\pi\hbar)^3} d\Omega p^2 dp =$$

$$= \frac{V}{(2\pi\hbar)^3} d\Omega \frac{E^2}{c^3} dE = \frac{V}{(2\pi\hbar)^3} \frac{\hbar^2 \omega^2}{c^3} dE$$

$$\rho(E) = \frac{V}{(2\pi\hbar)^3} \frac{p^2}{c} d\Omega = \quad dE = \hbar d\omega$$

$$\Rightarrow \underline{\hat{\rho}(\omega) = \frac{d^3n}{d\omega} = \frac{V}{(2\pi\hbar)^3} \left(\frac{\hbar}{c}\right)^3 \omega^2 d\Omega = \frac{V}{(2\pi c)^3} \omega^2 d\Omega}$$

How do we understand this in terms of Fermi-Golden rule?

$$R_{i \rightarrow f} \sim |\langle f | V | i \rangle|^2 \delta(E_f - E_i - \hbar\omega)$$

$\frac{P_{i \rightarrow f}}{t}$ prob. per unit time.

The point is: how many states do we have with

$$E_f - E_i = \hbar\omega? \quad R_{i \rightarrow f} = \sum_{\text{states}} \rightarrow \int dE \rho(E)$$

In case of more than one particle

$$\int \prod_K \frac{V d^3 p_K}{(2\pi\hbar)^3} \delta(E_f - E_i + \sum_K E_K) \delta(\vec{p}_f - \sum_K \vec{p}_K - \vec{p}_i) |\langle f | M | i \rangle|^2$$

$|f\rangle \rightarrow$ FINAL STATE SYSTEM

$\{K\} \rightarrow$ EMITTED PARTICLES

Example: decay $M \rightarrow$ 2 particle of mass m_1, m_2

$$\vec{p}_2, m_2 \longleftarrow \overset{\cdot}{M} \longrightarrow \vec{p}_1, m_1$$

$$\vec{p}_i = 0, E_i = Mc^2 \implies \vec{p}_f, E_f = 0 \text{ (no final state which is not a decay product)}$$

$$E_i = Mc^2 = \sqrt{(\vec{p}_1 c)^2 + m_1^2 c^4} + \sqrt{(\vec{p}_2 c)^2 + m_2^2 c^4}$$

$$\delta(\vec{p}_1 + \vec{p}_2 - \vec{0}) \implies \vec{p}_1 = -\vec{p}_2 \equiv \vec{p}$$

$$\implies \frac{V^2}{(2\pi\hbar)^6} \int d^3 \vec{p} \delta(Mc^2 - \sqrt{(\vec{p}c)^2 + m_1^2 c^4} - \sqrt{(\vec{p}c)^2 + m_2^2 c^4})$$

$$Mc^2 = \sqrt{p^2 c^2 + m_1^2 c^4} + \sqrt{p^2 c^2 + m_2^2 c^4}$$

$$M^2 c^4 = 2 p^2 c^2 + (m_1^2 + m_2^2) c^4 + 2 \sqrt{(p^2 c^2 + m_1^2 c^4)(p^2 c^2 + m_2^2 c^4)}$$

$$[M^2 c^4 - (m_1^2 + m_2^2) c^4] - 2 p^2 c^2 = 2 \sqrt{(\quad)(\quad)} \quad \text{✗}$$

Square again, call, $p^2 = x$

$$[\dots]^2 + 4x^2 c^4 - 4xc^2[\dots] = 4(xc^2 + m_1^2 c^4)(xc^2 + m_2^2 c^4)$$

$$\text{Solving for } p^2: p^2 = \frac{c^2}{4M^2} (M - m_1 - m_2)(M + m_1 - m_2)(M - m_1 + m_2)(M + m_1 + m_2)$$

$$p = \frac{c}{2M} \sqrt{(M - m_1 - m_2)(M + m_1 - m_2)(M - m_1 + m_2)(M + m_1 + m_2)} \quad \hookrightarrow \tilde{p}^2$$

$\hookrightarrow$ cannot decay if $m_1 + m_2 > M$.

$$\frac{V^2}{(2\pi\hbar)^6} \int d^3\tilde{p} \delta(Mc^2 - \sqrt{\tilde{p}^2 c^2 + m_1^2 c^4} - \sqrt{p^2 c^2 + m_2^2 c^4}) =$$

$$= \frac{V^2}{(2\pi\hbar)^6} 4\pi \cdot \int_0^\infty dp p^2 \delta(p - \tilde{p}) \cdot \underbrace{\frac{1}{\left| \frac{\partial}{\partial p} (Mc^2 - \sqrt{\quad} - \sqrt{\quad}) \right|}_{p=\tilde{p}}} =$$

$$= -\frac{1}{2} \frac{1}{\sqrt{m_1}} 2pc^2 - \frac{1}{2} \frac{1}{\sqrt{m_2}} 2pc^2 = -pc^2 \left(\frac{1}{\sqrt{m_1}} + \frac{1}{\sqrt{m_2}} \right)$$

$$= -pc^2 \left(\frac{\sqrt{\quad} + \sqrt{\quad}}{\sqrt{\quad}\sqrt{\quad}} \right) = \star$$

evaluating in $p=\tilde{p}$: numerator $= Mc^2$ (it must be)

denominator, using $\star$ $\frac{(Mc^4 - (m_1^2 + m_2^2)c^4) - 2p^2 c^2}{2} \Big|_{p=\tilde{p}}$

After some algebra:

$$\star = \frac{-c^3 \sqrt{(M-m_1-m_2)(M+m_1-m_2)(M-m_1+m_2)(M+m_1+m_2)} (M^2 - (m_1^2 - m_2^2)^2)}{M(M^2 - (m_1^2 - m_2^2)^2)}$$

But then there is still p^2 in numerator, solving last &

$$\Rightarrow = \frac{V^2}{(2\pi\hbar)^6} 4\pi \frac{(M^2 - (m_1^2 - m_2^2)^2)}{4cM(M^2 - (m_1^2 - m_2^2)^2)} \cdot \sqrt{(M-m_1-m_2)(M+m_1-m_2)(M-m_1+m_2)(M+m_1+m_2)}$$

2) Hydrogen Atom

At $t = -\infty \rightarrow$ ground state

An electric field is turned on at $t = -\infty$ and has the form

$$\vec{E}(t) = E \hat{n} e^{-t^2/\tau^2}$$

What is the probability of finding the atom in a state with $n=2$ at $t = +\infty$?

The time dependent potential is $V(t) = -e \vec{E} \cdot \vec{r}$

Choose $\hat{n} = \hat{e}_z \Rightarrow V(t) = -e z E e^{-t^2/\tau^2}$

$$C_{fi}^{(1)}(t) = -\frac{i}{\hbar} \int_{t_0}^t dt' e^{\frac{i}{\hbar}(E_f - E_i)t'} \langle f | V(t') | i \rangle =$$

$$= -\frac{i}{\hbar} \int_{t_0}^t dt' e^{\frac{i}{\hbar}(E_f - E_i)t'} e^{-t'^2/\tau^2} (-eE) \langle f | \hat{z} | i \rangle$$

$$P_{i \rightarrow f}(t \rightarrow \infty, t_0 \rightarrow -\infty) = \frac{e^2 E^2}{\hbar^2} \left| \int_{-\infty}^{+\infty} dt e^{-\left(\frac{t^2}{\tau^2} - i\omega_{fi}t\right)} \right|^2 |\langle f | \hat{z} | i \rangle|^2 \Rightarrow$$

$$\tau \int_{-\infty}^{+\infty} dt e^{-(t^2 - i\omega_{fi}\tau t)} = \tau e^{-\frac{\omega_{fi}^2 \tau^2}{4}} \int_{-\infty}^{+\infty} dt e^{-\left(t - \frac{i\omega_{fi}\tau}{2}\right)^2} =$$

$$= \tau \sqrt{\pi} e^{-\frac{\omega_{fi}^2 \tau^2}{4}}$$

$$\Rightarrow P_{i \rightarrow f}(t \rightarrow \infty, t_0 \rightarrow -\infty) = \frac{e^2 E^2}{\hbar^2} \tau^2 \pi e^{-\frac{\omega_{fi}^2 \tau^2}{2}}$$

We need to compute $\langle f | \hat{z} | i \rangle$ where

$$|i\rangle = |1,0,0\rangle \quad \langle r, \theta, \varphi | 1,0,0\rangle = \frac{1}{\sqrt{\pi} a_0^3} \exp\left(-\frac{r}{a_0}\right)$$

$$\langle n, \ell, m | \hat{z} | 1,0,0\rangle =$$

$$= \int_0^{+\infty} dr r^2 d\Omega \Psi_{n,\ell,m}^*(r, \theta, \varphi) r \cos\theta \frac{1}{\sqrt{\pi} a_0^3} \exp\left(-\frac{r}{a_0}\right)$$

$$= \int_0^{+\infty} dr r^2 d\Omega R_{n\ell}^*(r) Y_{\ell}^{m*}(\theta, \varphi) r \cos\theta \frac{1}{\sqrt{\pi} a_0^3} e^{-\frac{r}{a_0}}$$

Consider the angular part

$$\int_0^{2\pi} d\phi \int_{-1}^{+1} d\cos\theta Y_{\ell}^{m*}(\theta, \varphi) \cos\theta$$

$$\ell=0, m=0 \Rightarrow \int d\Omega \cos\theta = 0$$

$$\ell=1, \pm 1 \Rightarrow \int_0^{2\pi} d\phi e^{\pm i\phi} = 0$$

$$P_{100 \rightarrow 200} = 0$$

$$P_{100 \rightarrow 21\pm 1} = 0$$

$$\ell=1, m=0$$

$$\langle 2,1,0 | \hat{z} | 1,0,0\rangle = \frac{1}{\sqrt{\pi} a_0^3} \frac{1}{4} \frac{1}{\sqrt{2\pi} a_0^3} \int_0^{+\infty} dr r^2 \int_0^{2\pi} d\varphi \int_{-1}^{+1} d\cos\theta \cos^2\theta \frac{r}{a_0} e^{-\frac{r}{2a_0}} r \frac{1}{a_0}$$

$$= \frac{1}{\pi a_0^3 \sqrt{2}} 2\pi \frac{1}{3} \frac{1}{a_0} \int_0^{+\infty} dr r^4 e^{-\frac{3}{2} \frac{r}{a_0}} =$$

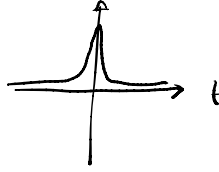
$$= \frac{1}{3\sqrt{2} a_0^3} a_0^4 \int_0^{+\infty} d\left(\frac{r}{a_0}\right) \left(\frac{r}{a_0}\right)^4 e^{-\frac{3}{2} \frac{r}{a_0}} = \frac{a_0}{3\sqrt{2}} \frac{256}{81} = \frac{128}{243} \sqrt{2} a_0 = \frac{2}{3^5} \sqrt{2} a_0$$

$$P_{200 \rightarrow 210} = \left(\frac{eE}{\hbar}\right)^2 \gamma_{\pi}^2 e^{-\frac{\omega_{af}^2 \gamma^2}{2}} \frac{2}{3^{10}} a_0^2$$

The total probability to jump in any $n=2$ state is then

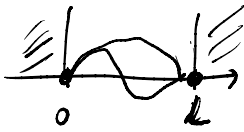
$$P_{gs \rightarrow n=2}(t \rightarrow \infty) = \left(\frac{eE}{\hbar}\right)^2 \left(\frac{2^{15}}{3^{10}} a_0^2\right) \pi \tau^2 e^{-\omega^2 \tau^2 / 2}$$

• For $\tau \rightarrow 0$ $P_{gs \rightarrow n=2} \rightarrow 0$



Sudden change

3)



$$\Psi(x) = N \sin(Kx) \quad K = \frac{n\pi}{L}, n \in \mathbb{Z}$$

$$\int_0^L dx N^2 \sin^2(Kx) = N^2 \frac{L}{2} \Rightarrow N = \sqrt{\frac{2}{L}}$$

$$\left\{ \begin{array}{l} \Psi_n(x) = \sqrt{\frac{2}{L}} \sin(Kx) \\ K = \frac{n\pi}{L}, n \in \mathbb{Z} \end{array} \right\}$$

$$E_n = \frac{\hbar^2}{2m} \frac{n^2 \pi^2}{L^2} \quad E_{gs} = \frac{\hbar^2 \pi^2}{2mL^2} = \frac{(mV)^2}{2m} = \frac{(m \frac{L}{T})^2}{2m}$$

$$\frac{\hbar^2 \pi^2}{L^2} = m^2 \frac{L^2}{T^2} \Rightarrow T^2 = \frac{m^2 L^4}{\hbar^2 \pi^2} \Rightarrow \boxed{T = \frac{mL^2}{\hbar \pi}}$$

The wave function for the ground state of the expanded box is:

$$\Psi_1^{2L}(x) = \sqrt{\frac{2}{2L}} \sin\left(\frac{\pi}{2L} x\right)$$

$$\langle gs_2 | gs_1 \rangle = \frac{\sqrt{2}}{L} \int_0^L dx \sin\left(\frac{\pi}{L} x\right) \sin\left(\frac{\pi}{2L} x\right)$$

$$P = |\langle \dots \rangle|^2 = \left| \frac{4}{3} \frac{L}{\pi} \frac{\sqrt{2}}{L} \right|^2 = \frac{16 \cdot 2}{9 \pi^2} = \frac{2^5}{(3\pi)^2}$$

4)

 H^3 (tritium)
 $\Psi_i = \Psi_{100}(r, \theta, \phi)$ initial state

$$E = \gamma m c^2 = 16 \text{ KeV} + 0.5 \text{ MeV} = 516 \text{ KeV}$$

$$\gamma \sim 1.0320 = \frac{1}{\sqrt{1 - v^2/c^2}} \Rightarrow v = c \cdot \sqrt{\frac{\gamma^2 - 1}{\gamma^2}} \approx 0.244 c$$

The atom has size $\sim a_0$.

$$v \sim \frac{a_0}{\tau_{\text{escape}}} \sim 7 \cdot 10^{-19} \text{ m}$$

Energy levels for single electron atoms $\sim (-13.6 \text{ eV}) \frac{Z^2}{n^2}$

$$E_0^{Z=1} = -13.6 \text{ eV} \quad \rightarrow Z=2$$

$$(-13.6 \text{ eV}) \left(-1 + \frac{4}{n^2} \right) = \Delta E = \left(\frac{4 - n^2}{n^2} \right) (-13.6 \text{ eV})$$

$$\tau_{\text{levels}} \sim \frac{n^2}{-4 + n^2} \frac{\hbar}{13.6 \text{ eV}} = \frac{n^2}{n^2 - 4} \quad 4.84 \cdot 10^{-17} \sim 10^{-17} \gg 10^{-19}$$

even for $n \gg 1$

sudden approx. can be used

The probability to fall into the ground state of Helium is

$$P = \left| \int d^3r \Psi_0^*(\vec{r}, Z=1) \Psi_0(\vec{r}, Z=2) \right|^2$$

$$= \left| \frac{\sqrt{8}}{\pi a_0^3} 4\pi \int_0^{+\infty} dr r^2 e^{-3r/a_0} \right|^2 =$$

$$= \frac{816}{200} a_0^6 \frac{4}{27^2} = \frac{512}{429} \sim 40\%$$

$$P |n=16, l=3, m=0\rangle = ? = 0.$$

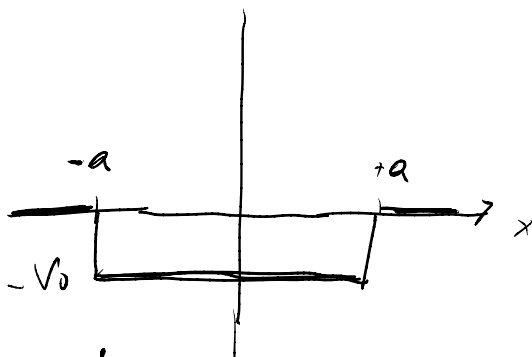
b) orthogonal

5)

$$V(x) = -V_0 \quad |x| < a$$

$p = \hbar k$ incoming

$$E \gg V_0$$



We quantise on a volume $L \gg a$

$$k = \frac{2\pi n}{L} \quad \text{and} \quad n \gg 1, \quad E \gg V_0$$

$$\psi_{\text{in}}(x) = \frac{1}{\sqrt{L}} e^{ikx}$$

$$\psi_{\text{out}}(x) = \frac{1}{\sqrt{L}} e^{-ikx}$$

To apply Fermi Golden Rule

$$\begin{aligned} \langle f | V(x) | i \rangle &= \int_{-L/2}^{L/2} dx \frac{1}{L} (e^{-ikx})^* V(x) e^{ikx} = \\ &= -V_0 \frac{1}{L} \int_{-a}^a dx e^{2ikx} = -\frac{V_0}{L} \frac{1}{2ik} (e^{2ika} - e^{-2ika}) \\ &= -\frac{V_0}{Lk} \sin(2ka) \Rightarrow R_{i \rightarrow f} = \frac{2\pi}{\hbar} V_0^2 \frac{\sin^2(2ka)}{L^2 k^2} g(E) \end{aligned}$$

Repeat exercise 1 in 1-d

$$dn = \frac{L}{2\pi\hbar} dp = \frac{L}{2\pi\hbar} \frac{m}{p} d\bar{e} \quad \left(\frac{p^2}{2m} = \bar{e} \quad \frac{dp}{m} = d\bar{e} \right)$$

$$\Rightarrow \rho(\bar{e}) = \frac{L}{2\pi\hbar} \frac{m}{p} = \frac{L}{2\pi\hbar} \frac{1}{2} \sqrt{\frac{2m}{\bar{e}}} = \frac{L}{4\pi\hbar} \sqrt{\frac{2m}{\bar{e}}}$$

$$R_{i \rightarrow f} = \frac{2\pi}{\hbar} V_0^2 \frac{\sin^2(2Ka)}{L^2 K^2} \frac{\sqrt{2m}}{4\pi\hbar} \sqrt{\frac{2m}{E}}$$

$$= \frac{V_0^2}{2\hbar^2} \sqrt{\frac{2m}{E}} \frac{\sin^2(2Ka)}{L K^2}$$

$$R_{i \rightarrow f} \xrightarrow{L \rightarrow \infty} 0$$

because the scattering potential is located in limited area $\ll L$, $L \rightarrow \infty$.

The space travelled in a time τ is:

$$v \cdot \tau_L = \sqrt{\frac{2}{m} E} \tau_L \stackrel{!}{=} L \Rightarrow \tau_L = \frac{L \sqrt{m}}{\sqrt{2E}}$$

$$\Rightarrow P_{\tau_L} = R_{i \rightarrow f} \tau_L = \frac{V_0^2}{2\hbar^2} \frac{\sqrt{2m}}{\sqrt{E}} \frac{\sqrt{m}}{\sqrt{2E}} \frac{\sin^2(2Ka)}{K^2}$$

$$= \frac{V_0^2}{2\hbar^2} \frac{m}{E} \frac{\sin^2(2Ka)}{K^2} = \frac{V_0^2}{4E^2} \sin^2(2Ka)$$

Solving the exact problem (in, out, normalisation, boundaries -- OK 1)



$$P_{refl} = \frac{(k/k - k/k)^2 \sin^2(2ka)}{(k/k + k/k)^2 \sin^2(2ka) + 4 \cos^2(2ka)}$$

where $k^2 - k^2 = 2mV_0/\hbar^2 \ll k^2$ so we can approximate, for $V_0 \ll E$, $ka \sim ka$ and $k/k \approx 1$

$$P_{refl} \approx \left(\frac{k^2 - k^2}{2k/k} \right)^2 \sin^2(2ka) \approx \frac{V_0^2}{4E^2} \sin^2(2ka)$$

Delta potential

The sudden approximation is based on the assumption that

$$\int_{-\varepsilon}^{+\varepsilon} dt H(t) |\psi(t)\rangle \xrightarrow{\varepsilon \rightarrow 0} 0$$

this is true for any "regular" function $H(t)$

But $\delta(t)$ has FINITE area over an INFINITESIMAL interval.

$\Rightarrow |\psi_0^+\rangle \neq |\psi_0^-\rangle$ in general.

Using perturbation theory:

$$c_{xf}(t) = -\frac{i}{\hbar} \int_{-\varepsilon}^{+\varepsilon} dt \underbrace{\langle f | H(t) | i \rangle}_{V \delta(t)} e^{i\omega_{fi}(t+\varepsilon)} = -\frac{i}{\hbar} \langle f | V | i \rangle$$