

9. Scattering theory: Generalities

SS 2024

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Scattering theory is the second MAIN TOPIC of this course.

Till now we have considered time-dependent processes as emission / absorption of light by atoms

⇒ SPECTROSCOPY important to understand Atoms & molecules

we concluded last lecture with the decay of $2p \rightarrow 1s$ of a Hydrogen Atom, and estimated

$$\text{a LIFE TIME } \tau \sim 1.6 \cdot 10^{-9} \text{ s}$$

We could see this as some sort of SCATTERING, some projectile hits Atom, excites it, till it decays

⇒ IMPORTANT POINT interaction time-scale MUCH SHORTER than "life-time" of Atom

$$\tau_{\text{characteristic}} \sim \frac{a_0}{c} \sim 2 \cdot 10^{-17} \text{ s}$$

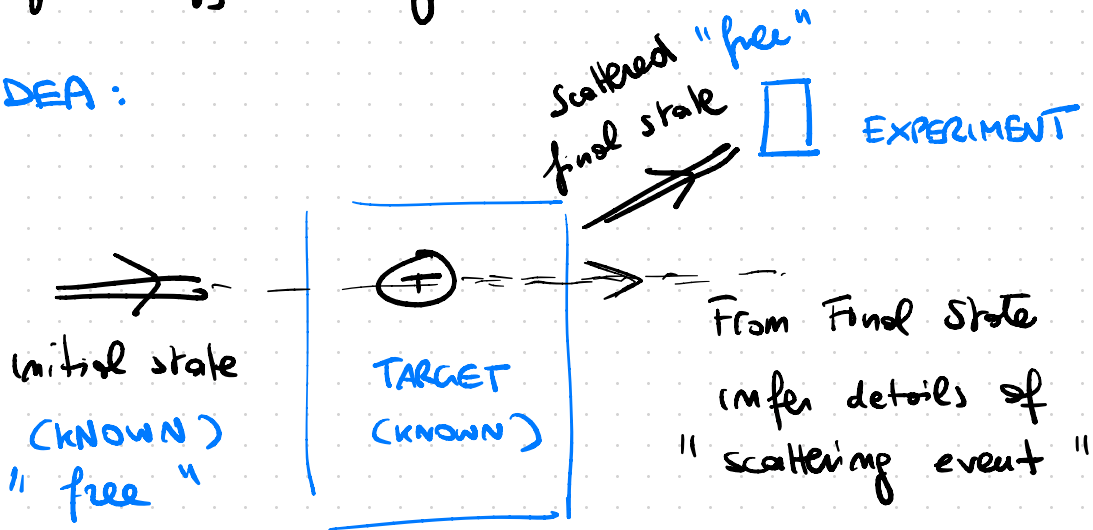
time for e^-
to revolve
around nucleus

So we can imagine to "separate" DECAY from EXCITATION such that details of how state decays are independent from the actual process that generated excited state

In SCATTERING [COLLISION] THEORY we don't make this assumption anymore
⇒ We discuss the PROCESS at ONCE as a WHOLE

Crucial to understand NUCLEAR STRUCTURE, high-energy scattering (CERN LHC etc)

IDEA:



In back of our minds, we imagine to work with **WAVE PACKETS** \Rightarrow $\begin{cases} \text{Larger than target} \\ \text{Smaller than ~~target~~ } \end{cases}$

In practice, we simplify our treatment working with **PLANE WAVES** \Rightarrow solutions of FREE

SCHRODINGER EQUATION

Let us recap some details

$$i\hbar \frac{\partial \psi(\vec{x}, t)}{\partial t} = H \psi(\vec{x}, t) \quad \text{3 Dim Schröd. Eq.}$$

$$H = \frac{\vec{p}^2}{2m} + V(\vec{x}) \quad V(\vec{x}) \in \mathbb{R}$$

Remember that PROBABILITY DENSITY $\rho = \psi^* \psi$:

$$\begin{aligned} \frac{\partial}{\partial t}(\psi^* \psi) &= \psi^* \frac{1}{i\hbar} \left[-\frac{\hbar^2}{2m} \vec{\nabla}^2 \psi + V\psi \right] \\ &\quad - \frac{1}{i\hbar} \left[-\frac{\hbar^2}{2m} \vec{\nabla}^2 \psi^* + V\psi^* \right] \psi = \end{aligned}$$

$$= \frac{i\hbar}{2m} \left[\psi^* \vec{\nabla}^2 \psi - \vec{\nabla}^2 (\psi^*) \psi \right]$$

$$= \vec{\nabla} \cdot \left[\frac{i\hbar}{2m} \left(\psi^* \vec{\nabla} \psi - (\vec{\nabla} \psi^*) \psi \right) \right]$$

$$\underbrace{\vec{j} = -\frac{i\hbar}{2m} (\psi^* \vec{\nabla} \psi - \psi \vec{\nabla} \psi^*)}_{\text{CURRENT}} \Rightarrow \underbrace{\frac{\partial \rho}{\partial t} = -\vec{\nabla} \cdot \vec{j}}_{\text{CURRENT CONSERVATION}}$$

if PLANE WAVE $\psi = A e^{-i(Et - \vec{p} \cdot \vec{x})/\hbar}$

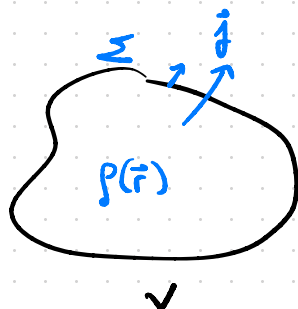
$$\begin{aligned} \vec{j} &= |A|^2 \frac{\vec{p}}{m} \propto \text{velocity of wave!} \\ &= |A|^2 \frac{\hbar \vec{k}}{m} \end{aligned}$$

\vec{j} current also called FLUX \Rightarrow # of particles per unit time crossing a unit area normal to incident direction

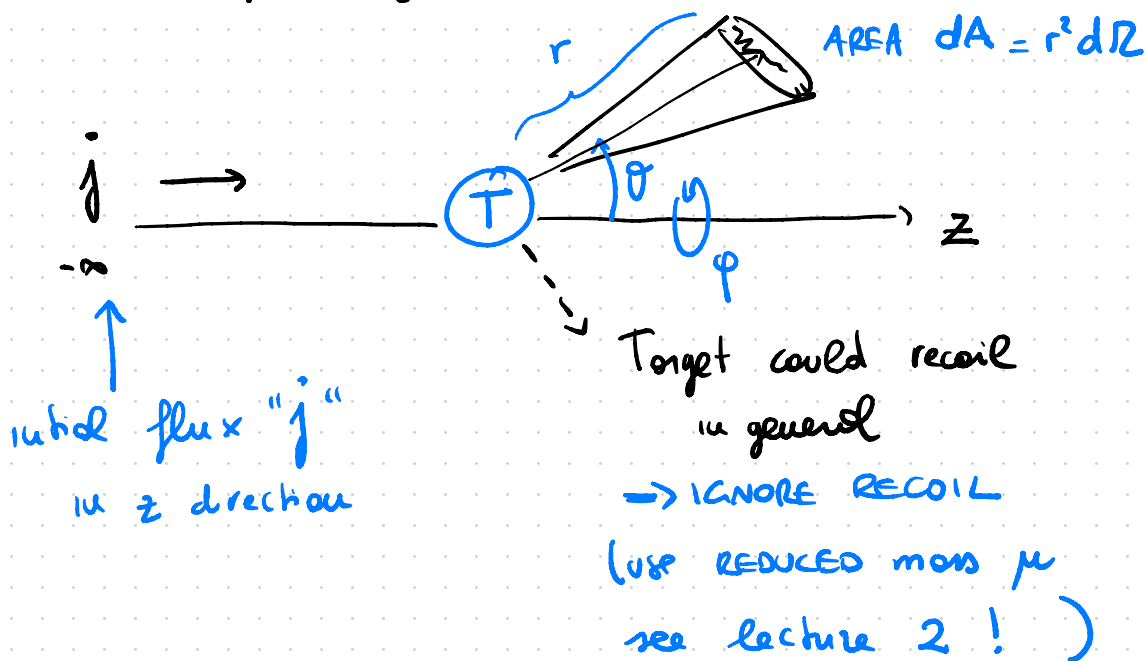
$$\Rightarrow \int_V \frac{\partial P}{\partial t} d^3\vec{r} = - \int \vec{\nabla} \cdot \vec{j} d^3\vec{r}$$

$$\frac{\partial}{\partial t} [P] = - \int_{\Sigma} \vec{j} \cdot d\vec{\sigma}$$

Probability to cross Σ
in unit time



Consider following set-up



θ POLAR ANGLE

ϕ AZIMUTHAL ANGLE

determine direction
of scattered particle

if the detector measures δn particles / second, then

$$\delta n = j \left[\frac{d\sigma(\theta, \varphi)}{d\Omega} \right] d\Omega$$

\uparrow RATE
 \uparrow INCOMING FLUX
 \uparrow DIFFERENTIAL CROSS-SECTION (per solid angle $d\Omega$)

AREA IS $r^2 d\Omega = dA$
 (notice no dA here!)

$$\frac{\delta n}{j dA} \left\{ \begin{array}{l} \text{measurable} \\ \text{quantities} \end{array} \right\} \Rightarrow \frac{d\sigma}{d\Omega}$$

can be computed in Q.M.

DEPENDS ON POTENTIAL V

We can then turn problem around and use measured $d\sigma$ to INFER POTENTIAL V !

Start from Schrödinger Equation

$$\left[-\frac{\hbar^2}{2\mu} \nabla_{\vec{x}}^2 + V(\vec{x}) \right] \psi = E \psi$$

• if $E < 0$ bound state in $V(x)$

\Rightarrow Hydrogen Atom, no Scattering !

• if $E > 0$ continuum of solutions for E

Define $E = \frac{\hbar^2}{2\mu} k^2$; $V(x) = \frac{\hbar^2}{2\mu} U(x)$

so Schrödinger Equation becomes

$$\Rightarrow (\vec{\nabla}^2 + k^2) \psi - U(x) \psi = 0$$

SCATTERING EQUATION

Now assuming interaction happens in SMALL REGION where $V \neq 0$, we start looking at form of solution ASYMPTOTICALLY FAR AWAY where $V = 0$!

Start with INCIDENT PLANE WAVE along z

$$E = \frac{\hbar^2}{2\mu} k^2$$

$$\psi_{\text{inc}} = \underset{\substack{\uparrow \\ \text{normalization}}}{1} e^{ikz}$$

$$\Rightarrow (\nabla^2 + k^2) \psi_{\text{inc}} = 0 \quad \text{of course}$$

[note of this ψ_{inc} continues unscattered!]

After interaction $\psi_{\text{scat}} \Rightarrow$ particles come out
"in any direction", solution should somehow
still solve FREE SCATTERING EQUATION

Let's write it in SPHERICAL COORDINATES

$$(\nabla^2 + k^2) \psi = \left[\frac{1}{r} \frac{\partial^2}{\partial r^2} (r\psi) + k^2 \psi \right]$$

$$+ \frac{1}{r^2} \left[\frac{1}{\sin\theta} \frac{\partial}{\partial \theta} \left(\sin\theta \frac{\partial \psi}{\partial \theta} \right) + \frac{1}{\sin^2\theta} \frac{\partial^2 \psi}{\partial \phi^2} \right]$$
$$= 0 \quad - \frac{L^2}{\hbar^2}$$

as we are interested in asymptotic solution ($r \rightarrow \infty$)
we try the following Ansatz

$$\psi_{\text{scat}} = e^{ikr} \left[\sum_{i=1}^{\infty} f_k^{(i)}(\vartheta, \varphi) \left(\frac{1}{r}\right)^i \right] \Rightarrow (kr = |\vec{k}| |\vec{r}|)$$

$$= \frac{e^{ikr}}{r} \left[f_k(\vartheta, \varphi) + \mathcal{O}\left(\frac{1}{r}\right) \right]$$

$$\frac{1}{r} \frac{\partial^2}{\partial r^2} \left[\cancel{r} \frac{e^{ikr}}{\cancel{r}} f_k(\vartheta, \varphi) \right] = -k^2 \frac{e^{ikr}}{r} f_k(\vartheta, \varphi)$$

$$\frac{1}{r} \frac{\partial^2}{\partial r^2} \left[\frac{e^{ikr}}{r} \frac{g(\vartheta, \varphi)}{r} \right] \propto -k^2 \frac{e^{ikr}}{r} \frac{g(\vartheta, \varphi)}{r} + \mathcal{O}\left(\frac{1}{r^3}\right)$$

↑
higher order
corrections

$$\frac{1}{r^2} \left[-\frac{L^2(\vartheta, \varphi)}{\hbar^2} \right] \left(\frac{e^{ikr}}{r} f_k(\vartheta, \varphi) + \mathcal{O}\left(\frac{1}{r}\right) \right) \propto \mathcal{O}\left(\frac{1}{r^2}\right)$$

$\Rightarrow \psi_{\text{scat}}$ solves scattering equation asymptotically! q

$f_k(\theta, \varphi)$ is a function of the angles only and
 the momentum is unconstrained

$f_k(\theta, \varphi)$ called SCATTERING AMPLITUDE

A full solution asymptotically $[V(\vec{x}) = 0]$ is

$$\psi = e^{ikz} + \frac{e^{ikr}}{r} f_k(\theta, \varphi) [1 + O(\frac{1}{r})]$$

Now let's compute the current \vec{j} ($e^{ikz} = e^{i\vec{k} \cdot \vec{r}}$)

$$\vec{j} = -\frac{i\hbar}{2\mu} (\psi^* \vec{\nabla} \psi - \psi \vec{\nabla} \psi^*)$$

$$= \frac{\hbar}{2i\mu} \left\{ \left[e^{-ikz} + f_k^*(\theta, \varphi) \frac{e^{-ikr}}{r} \right] \times \right.$$

$$\times \left[ik e^{ikz} \hat{z} + \left(ik \frac{e^{ikr}}{r} - \frac{e^{ikr}}{r^2} \right) f_k(\theta, \varphi) \hat{r} + \frac{1}{r} \frac{\partial f_k}{\partial \theta} \frac{e^{ikr}}{r} \hat{\theta} + \frac{1}{r \sin \theta} \frac{\partial f_k}{\partial \varphi} \hat{\varphi} \right]$$

- complex conjugated }

evaluating the products we get

$$\vec{j} = \frac{\hbar k}{\mu} \hat{z} + \frac{\hbar k}{\mu} \frac{\hat{r}}{r^2} |f_k(\vartheta, \varphi)|^2 + O\left(\frac{1}{r^3}\right)$$

+ interference terms with e^{ikr} & e^{-ikz} !

• $O\left(\frac{1}{r^3}\right)$ clearly suppressed at large distance !

• What about INTERFERENCE TERMS ?
they contain both types of exponentials

$$j_{\text{int}} \propto A e^{ikr} e^{-ikz} + \text{c.c.}$$

$$\sim A e^{ikr(1-\cos\vartheta)} + \text{c.c.} \quad \left(\text{using } \underline{z = r \cos\vartheta} \right)$$

NOTICE : we always care about $\vartheta \neq 0$ (there is no way to distinguish scattered from unscattered particles in FORWARD DIRECTION!) 11

Now in a **REALISTIC EXPERIMENT** we must integrate \vec{j} over some small $d\Omega$

$$\sim \int d\omega d\varphi g(\vartheta, \varphi) e^{ikr(1-\cos\vartheta)} A$$

↑
unspecified smooth "acceptance" function for the DETECTOR

when $r \rightarrow \infty$, integral of some smooth function times a RAPIDLY OSCILLATING ONE $e^{ikr(1-\cos\vartheta)}$

Riemann - Lebesgue lemma says this

integral VANISHES FASTER THAN ANY $\frac{1}{r^n}$!
(SEE EXERCISES!)

so we can neglect all interference terms and

$$\boxed{\vec{j} = \frac{\hbar k}{\mu} \hat{z} + \frac{\hbar k}{\mu} \frac{\hat{r}}{r^2} |f_k(\vartheta, \varphi)|^2}$$

TOTAL
FLUX
FAR AWAY!

Using this current in our formula for the cross-section gives

$$d\sigma(\vartheta, \varphi) = \underbrace{\int_{\mathbf{I}}^{-1}}_{\substack{\text{INCIDENT} \\ \text{FLUX} \sim \frac{1}{2} \frac{\hbar k}{\mu}}} \cdot \underbrace{\int_{\text{SCATT}}}_{\substack{\text{SCATTERING RATE} \Rightarrow \text{FLUX SCATTERED!} \\ \sim \frac{\hat{r}}{r^2} \frac{\hbar k}{\mu} |f_k(\vartheta, \varphi)|^2}} \cdot \underbrace{dA}_{\substack{\text{AREA SUBTENDS} \\ \text{SOLID ANGLE } d\Omega}} = \frac{\mu}{\hbar k} \left[\frac{1}{r^2} \frac{\hbar k}{\mu} |f_k(\vartheta, \varphi)|^2 \right] r^2 d\Omega$$

$$\Rightarrow \boxed{d\sigma(\vartheta, \varphi) = |f_k(\vartheta, \varphi)|^2 d\Omega} \quad \text{CROSS SECTION}$$

Notice that terms $\propto O(\frac{1}{r^3})$ in \vec{f} would contribute to non-section with $\propto O(\frac{1}{r})$ due to the r^2 at the numerator of integration measures \Rightarrow confirms it is ok to neglect them! B

more explicitly: our starting formula was

$$\delta n = \int_{\mathbb{I}} \left(\frac{d\sigma(\vartheta, \varphi)}{d\Omega} \right) d\Omega \quad \text{which implies}$$

$$\frac{d\sigma}{d\Omega} = \int_{\mathbb{I}}^{-1} \frac{\delta n}{d\Omega}$$

now δn is rate
measured by detector

$$\Rightarrow \delta n = \int_{\text{surf}} \times dA$$

$dA = r^2 d\Omega$
area subtended
by $d\Omega$ at detector!

$$\Rightarrow \frac{d\sigma}{d\Omega} = \int_{\mathbb{I}}^{-1} \frac{\int_{\text{surf}} \cdot r^2 \cdot d\Omega}{d\Omega}$$

$$= \frac{\mu}{\hbar k} \cdot \frac{1}{r^2} \cdot \frac{\hbar \mu}{\mu} |f_n(\vartheta, \varphi)|^2 \cdot \frac{r^2 d\Omega}{d\Omega}$$

$$= |f_n(\vartheta, \varphi)|^2$$

\Rightarrow in order to compute cross-section, we need
the scattering amplitude

To compute it, we need to solve Schrödinger Eq.
and see what information is brought in by the
potential $V(x)$. To do that, we use the
GREEN'S FUNCTION method, that you should know
from Electrodynamics.

Start from $(\underbrace{\vec{\nabla}^2 + k^2})\psi = 0$

Free Schrödinger Operator

Green's function $G(x)$ is defined as

$$(\vec{\nabla}^2 + k^2) G(\vec{x}) = \delta^{(3)}(\vec{x}) \leftarrow \text{Dirac } \delta \text{ function in 3-dim}$$

$$\delta^{(3)}(\vec{x}) = \delta(x_1) \delta(x_2) \delta(x_3) \quad \begin{aligned} \vec{x} &= (x_1, x_2, x_3) \\ &\sim (x, y, z) \end{aligned}$$

indeed, if we have $G(\vec{x})$, then we can write

$$\boxed{\psi(\vec{x}) = \psi_0(\vec{x}) + \int d^3\vec{y} G(\vec{x}-\vec{y}) U(\vec{y}) \psi(\vec{y})}$$

with $(\vec{\nabla}^2 + k^2) \psi_0(\vec{x}) = 0$

then acting on $\psi(\vec{x})$ with $\vec{\nabla}^2 + k^2$ we get

$$\begin{aligned} (\vec{\nabla}_{\vec{x}}^2 + k^2) \psi &= 0 + \int d^3\vec{y} (\vec{\nabla}_{\vec{x}}^2 + k^2) G(\vec{x}-\vec{y}) U(\vec{y}) \psi(\vec{y}) \\ &= \int d^3\vec{y} \delta^{(3)}(\vec{x}-\vec{y}) U(\vec{y}) \psi(\vec{y}) \\ &= U(\vec{x}) \psi(\vec{x}) \end{aligned}$$

The boxed equation above is a solution of the

Schrödinger Equation \Rightarrow we still need to make sure

we fix BOUNDARY CONDITION to get physical solution!

$\psi_0(x)$ & $G(\vec{x}-\vec{x}')$

Before thinking about how to compute $G(\vec{x})$,
Notice that this equation is sorta an integral
Equation, very similar to the one we wrote
when studying **time-dependent perturbation theory**.

It becomes **USEFUL** if we assume that $U(\vec{y})$ is
"small" and we can then **ITERATE** the
equation to get a **Series Expansion** in U
 \Rightarrow **"BORN" EXPANSION (SERIES)**

$$\begin{aligned}\psi(x) = & \psi_0(x) + \int d^3\vec{y} \, G(\vec{x}-\vec{y}) \, U(\vec{y}) \, \psi_0(\vec{y}) \\ & + \int d^3\vec{y} \, G(\vec{x}-\vec{y}) \, U(\vec{y}) \int d^3\vec{z} \, G(\vec{y}-\vec{z}) \, U(\vec{z}) \, \psi_0(\vec{z}) \\ & + O(U(x)^3) \quad \text{etc}\end{aligned}$$

\Rightarrow useful because $\psi_0(\vec{x})$ is known!

$$\psi_0(x) = e^{ikz} \quad \text{incoming (unscattered) wave}$$

the first order is called

(first) BORN APPROXIMATION

$$\psi_{\text{scat}}^{(1)}(\vec{x}) = \frac{2\mu}{\hbar^2} \int d^3\vec{y} \, G(\vec{x}-\vec{y}) \, V(\vec{y}) \, e^{i\vec{k}\cdot\vec{y}}$$

back to
STANDARD POTENTIAL

incoming wave

$$\{ \vec{k} = (0, 0, k) ; \vec{y} = (y_1, y_2, y_3) \}$$

To make use of this formula, we need $G(\vec{x}-\vec{y})$!

We will prove more properly in next lecture that

$$G(\vec{x}) = G(|\vec{x}|) = - \frac{e^{+i|\vec{x}|}}{4\pi|\vec{x}|}$$

is right solution for OUTGOING SPHERICAL WAVE

To prove it, we should see that :

$$(\vec{\nabla}^2 + k^2) G(|\vec{x}|) = (\vec{\nabla}^2 + k^2) G(r) \stackrel{?}{=} \delta^{(3)}(\vec{x})$$

\uparrow
spherical coordinates

as long as $r \neq 0$ we can just differentiate :

$$= \frac{1}{r} \frac{\partial^2}{\partial r^2} (r G(r)) + k^2 G(r)$$

$$= -\frac{1}{r} \frac{\partial^2}{\partial r^2} \left[\cancel{r} \frac{e^{+ikr}}{4\pi \cancel{r}} \right] - k^2 \frac{e^{+ikr}}{4\pi r}$$

$$= + \frac{k^2}{r 4\pi} e^{+ikr} - k^2 \frac{e^{+ikr}}{4\pi r} = 0 \quad \checkmark$$

$\Rightarrow @ r=0$ $\frac{\partial}{\partial r}$ etc all well defined

We should find $\delta^{(3)}(\vec{x}) \Rightarrow$ DISTRIBUTION,
makes sense only upon integration on some
region of space \Rightarrow SOME SPHERE CENTRED @ $\vec{x}=0$

$$\int_{S_\epsilon} dV \vec{\nabla}^2 \psi(r) =$$

S_ϵ sphere of radius ϵ

$$\stackrel{\text{DIVERGENCE THEOREM}}{=} = \int_{S_\epsilon} (\vec{\nabla} \cdot \psi) \cdot \hat{r} d\sigma$$



now $\vec{\nabla} \cdot \psi(r) = \hat{r} \frac{\partial \psi}{\partial r}$

$$= \frac{\hat{r}}{4\pi} \left[\frac{1}{r^2} - \frac{ik}{r} \right] e^{ikr}$$

$$= \int_{S_\epsilon} \frac{1}{4\pi} \left[\frac{1}{r^2} - \frac{ik}{r} \right] e^{ikr} \underset{\uparrow}{d\sigma}$$

surface is @ r
fixed $= \epsilon$

$$\equiv \cancel{4\pi} \epsilon^2 \frac{1}{\cancel{4\pi}} \left[\frac{1}{\epsilon^2} - \frac{ik}{\epsilon} \right] e^{ik\epsilon}$$

$$= \underline{1 + O(\epsilon^2)}$$

so we find that as $\varepsilon \rightarrow 0$, $\int_{S_\varepsilon} \nabla^2 G dV = 1$

which is exactly what we expect from $\delta^{(3)}(\vec{x})$

\Rightarrow for other smooth functions integral would go to zero when volume shrinks

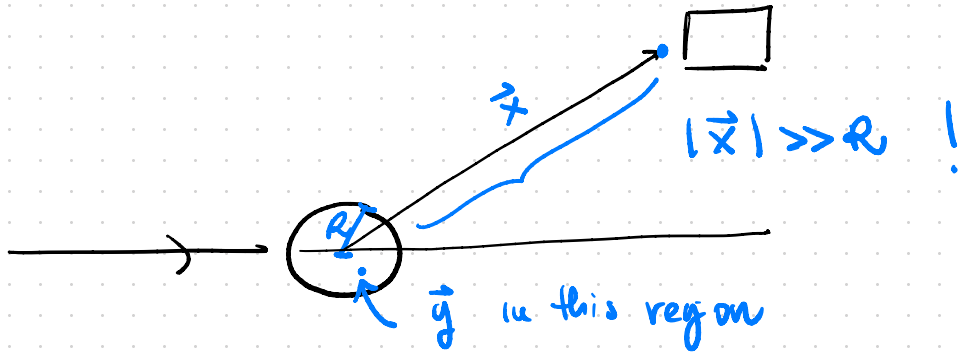
Now, using $G(\vec{x}) = -\frac{e^{+ik|\vec{x}|}}{4\pi|\vec{x}|}$ in formula for

FIRST ORDER Born Approximation

$$\begin{aligned}\psi_{\text{scat}}^{(1)}(\vec{x}) &= \frac{2\mu}{\hbar^2} \int d^3\vec{y} \, G(\vec{x}-\vec{y}) \, V(\vec{y}) \, e^{ik y_3} \\ &= -\frac{\mu}{2\pi\hbar^2} \int d^3\vec{y} \, \frac{e^{+ik|\vec{x}-\vec{y}|}}{|\vec{x}-\vec{y}|} \, V(\vec{y}) \, e^{ik y_3}\end{aligned}$$

(remember $y_3 = \hat{z} \cdot \vec{y}$; etc)

Typically $V(\vec{x}) \neq 0$ in some bounded region $|\vec{x}| < R \Rightarrow$ FINITE RANGE POTENTIAL



then we can simplify the integral using

$$|\vec{x} - \vec{y}| = \sqrt{\vec{x}^2 + \vec{y}^2 - 2\vec{x} \cdot \vec{y}} = |\vec{x}| \sqrt{1 - 2 \frac{\hat{x} \cdot \vec{y}}{|\vec{x}|} + \frac{|\vec{y}|^2}{|\vec{x}|^2}}$$

$$\approx |\vec{x}| - \hat{x} \cdot \vec{y} + O\left(\frac{|\vec{y}|^2}{|\vec{x}|^2}\right) \quad \text{and}$$

$$\psi_{\text{scat}}^{(1)}(\vec{x}) \approx - \frac{\mu e^{ik|\vec{x}|}}{2\pi \hbar^2 |\vec{x}|} \int d^3\vec{y} e^{-ik\hat{x} \cdot \vec{y}} e^{ik\hat{z} \cdot \vec{y}} V(\vec{y})$$

$$= - \frac{\mu}{2\pi\hbar^2} \frac{e^{ik|\vec{x}|}}{|\vec{x}|} \int d^3\vec{y} e^{i(\vec{k}_i - \vec{k}_f) \cdot \vec{y}} V(\vec{y})$$

where $\begin{cases} \vec{k}_i = (0, 0, k) \\ \vec{k}_f = k \hat{x} \end{cases}$ wave vector of incident particle
 wave vector of scattered particle

notice that this is exactly in form

$$\psi = e^{ikz} + \frac{e^{ikr}}{r} f_k(\theta, \varphi) [1 + o(\frac{1}{r})]$$

with

$$f_k(\theta, \varphi) = - \frac{\mu}{2\pi\hbar^2} \int d^3\vec{y} e^{i(\vec{k}_i - \vec{k}_f) \cdot \vec{y}} V(\vec{y})$$

\Rightarrow the SCATTERING AMPLITUDE in BORN APPROX
 is given by the FOURIER TRANSFORM
 of the SCATTERING POTENTIAL

