Relativity, Particles, Fields SS 2017

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Sheet 6: Classical Source, Wick's Theorem (13.6.2017)

1 Particle Creation by a Classical Source

In this exercise we return to the particle creation by a classical source. Recall from Sec. 5.8.2 of the script that this process can be described by the Hamiltonian

$$H = H_0 - \int d^3x J(x)\phi(x), \qquad (1)$$

where H_0 is the free Klein-Gordon Hamiltonian, $\phi(x)$ is the Klein-Gordon field, and J(x) is the classical, external source (a c-number scalar function). We found that, if the system is in the vacuum state before the source is turned on, the source will create a mean number of particles

$$\langle N \rangle = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{2E_p} |\widetilde{J}(p)|^2 \,, \tag{2}$$

where \widetilde{J} is the Fourier transform of J. In this problem we will verify that statement, and extract more detailed information, by using a perturbative expansion in the strength of the source.

a) Show that the probability that the source creates *no* particles is given by

$$P(0) = \left| \langle 0|T \left\{ \exp\left[i \int d^4 x J(x) \phi_I(x) \right] \right\} |0\rangle \right|^2.$$
(3)

b) Evaluate the term in P(0) of order J^2 , and show that $P(0) = 1 - \lambda + O(J^4)$, where $\lambda = \langle N \rangle$ equals the expression given in Eq. (2).

c) Represent the term computed in b) as a Feynman diagram. Now represent the whole perturbation series for P(0) in terms of Feynman diagrams. Show that this series exponentiates, so that it can be summed exactly: $P(0) = e^{-\lambda}$.

Hint: the term computed in b) can be represented as a propagator, and therefore the whole perturbation series for P(0) can be written as an infinite sum over non-interacting propagators, with appropriate symmetry factors.

d) Working at order J^2 , compute the probability that the source creates one particle of momentum **k**. Then integrate this probability over **k** to show that the probability to create one particle with *any* momentum is, at $O(J^2)$, $P(1) = \lambda$. Finally, using a representation similar to that employed in point c), show that the all-orders expression of this probability is $P(1) = \lambda e^{-\lambda}$.

e) The all-orders results obtained above for n = 0 and n = 1 can be generalized to show that the probability to create n particles is

$$P(n) = \frac{\lambda^n e^{-\lambda}}{n!},\tag{4}$$

i.e. it follows the Poisson distribution. Accepting this fact, prove the following properties of P(n):

$$\sum_{n=0}^{\infty} P(n) = 1, \qquad \langle n \rangle = \sum_{n=0}^{\infty} n P(n) = \lambda.$$
(5)



The first property ensures that the P(n) are correctly normalized probabilities, while the second confirms the interpretation of Eq. (2) as the mean number of particles created by the source J.

2 Applications of Wick's Theorem

Working in the $\lambda \phi^4$ theory,

$$\mathcal{L} = \frac{1}{2} \partial_{\mu} \phi \partial^{\mu} \phi - \frac{1}{2} m^2 \phi^2 - \frac{\lambda}{4!} \phi^4$$

(all expressions are in the interaction picture, $\phi(x) = \phi_I(x)$), use Wick's theorem to evaluate the following quantities:

- **a)** $\langle 0|T\{\phi(x_1)\phi(x_2)\phi(x_3)\phi(x_4)\}|0\rangle$,
- **b)** $\langle 0|T\{\phi(x_1)\phi(x_2)\int d^4x_3 \frac{\lambda}{4!}\phi^4(x_3)\}|0\rangle$,
- c) $\langle 0|T\{\phi^4(x_1)\phi^4(x_2)\}|0\rangle$,

by expressing them in terms of Feynman propagators. In each case, also draw the Feynman diagrams representing the result.