## GENERAL RELATIVITY WS 2017/2018 Technische Universität München November 20, 2017

## Exercise Sheet $5^*$

The solutions to the following problem set should be handed in by the 24th of November at 8:30 a.m. at the postbox next to PH 3218.

- 1. Prove that all geodesics on  $S^2$  are great circles (meaning circles which are contained in a plane that passes through the origin, e.g. the equator).
- 2. Given a diagonal metric  $g_{\mu\nu}$ , prove that the Christoffels symbols, can be computed in the following way, assume  $\mu \neq \nu \neq \lambda$  and that there is no summation convention:

$$\Gamma^{\lambda}_{\mu\nu} = 0 \tag{1}$$

$$\Gamma^{\lambda}_{\mu\mu} = -\frac{1}{2g_{\lambda\lambda}}\partial_{\lambda}g_{\mu\mu} \tag{2}$$

$$\Gamma^{\lambda}_{\mu\lambda} = \partial_{\mu} \left( \ln \sqrt{|g_{\lambda\lambda}|} \right) \tag{3}$$

$$\Gamma^{\lambda}_{\lambda\lambda} = \partial_{\lambda} \left( \ln \sqrt{|g_{\lambda\lambda}|} \right) \tag{4}$$

3. The metric close to the Earth's surface can be accurately described by:

$$ds^{2} = -(1+2\Phi)dt^{2} + (1-2\Phi)dr^{2} + r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2})$$
(5)

with

$$\Phi = -\frac{GM}{r} \tag{6}$$

where G is Newton's constant and M is the Earth's mass.

- (a) Imagine two clocks at distances  $R_1$  and  $R_2$  from the center of the Earth, with  $R_2 > R_1$ . Which clock ticks faster?
- (b) Solve the geodesic equation for an orbit constrained to the equator,  $\theta = \pi/2$ . What is the angular velocity  $d\phi/dt$ ?
- (c) During one orbit of such geodesic at a radius R, how much proper time elapses? (assume  $\Phi \ll 1$ ) Make R equal to the radius of the earth and get an answer in seconds. Compare this number with the proper time elapsed for the stationary clock in item 3a.
- 4. Generalize the result for the curvature of a sphere seen in the lecture, to the *n*-dimensional case.

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